New Views of Planum Boreum Interior in a Migrated 3D Volume of SHARAD Data

SHARAD: The Shallow Radar on the Mars Reconnaissance Orbiter

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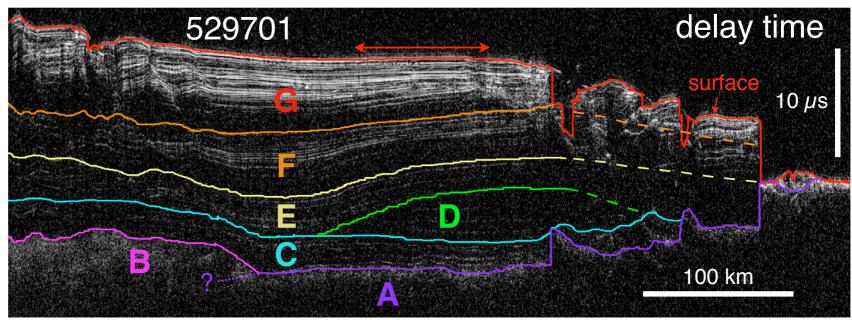
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 ²Freestyle Analytical & Quantitative Services, LLC, Longmont, Colorado
 ³National Air & Space Museum, Smithsonian Institution, Washington, D.C.

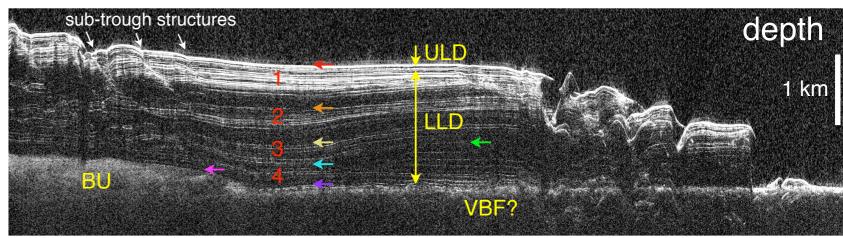
8th International Conference on Mars

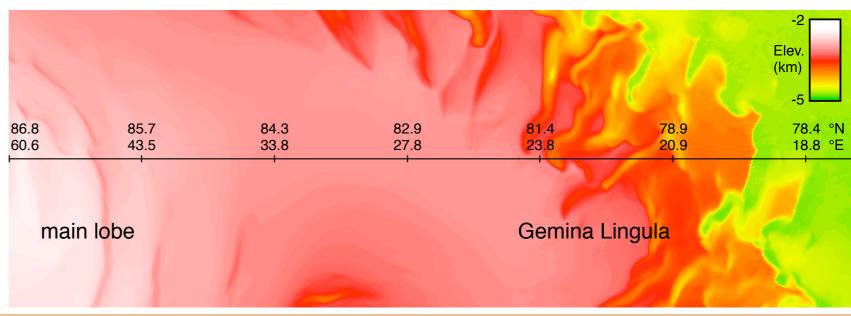
Image credit: NASA/JPL

2D SHARAD analysis

- Delineate units and reflectors.
- Interpolate
 through clutter
 and between
 orbital tracks.
- Map reflecting surfaces in 3D, calculate volumes.
- Drill, baby, drill!





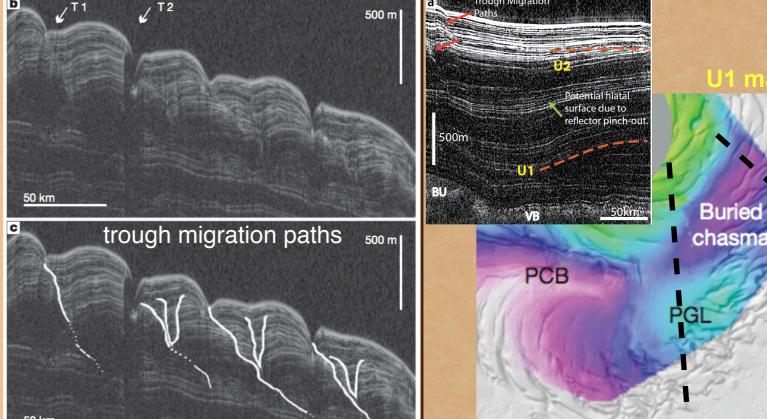


Putzig et al. (Icarus 2009)

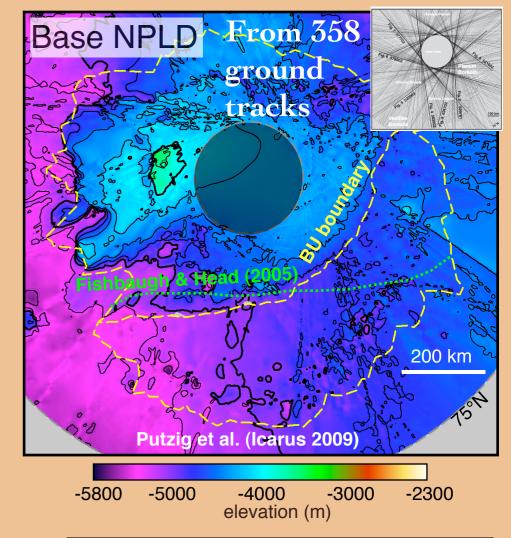
2D SHARAD analysis is yielding great results!

- Revised BU boundary rewrites Chasma Boreale history—it's constructional, not erosional.
- Spiral troughs proven to be wind-driven icy bedforms.*
- Buried chasma with no surface
 expression is
 revealed!

* See next talk by Isaac Smith



Smith and Holt (Nature 2010)

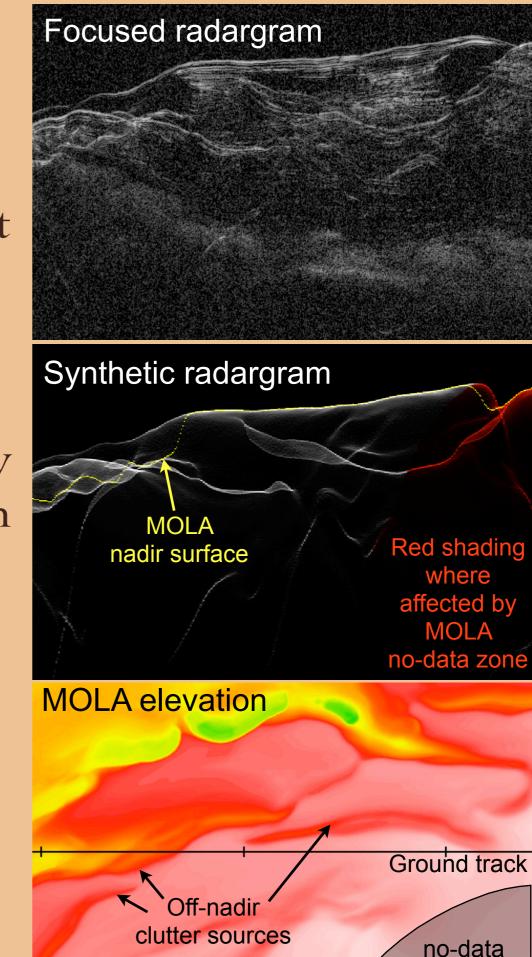


Holt et al. (Nature 2010)

Limitations of 2D SHARAD analysis

- No returns from features at nadir that slope away from the radar.
- Off-nadir returns (clutter signals) interfere with or are mistaken for nadir returns. Synthetics help identify surface-clutter signals, which are then dismissed as "noise".
- In many areas, intense clutter makes the data largely uninterpretable.

Mapping features at depth is often challenging and tedious.



zone

Inherent problem: 2D sampling of a 3D signal returned from a 3D structure

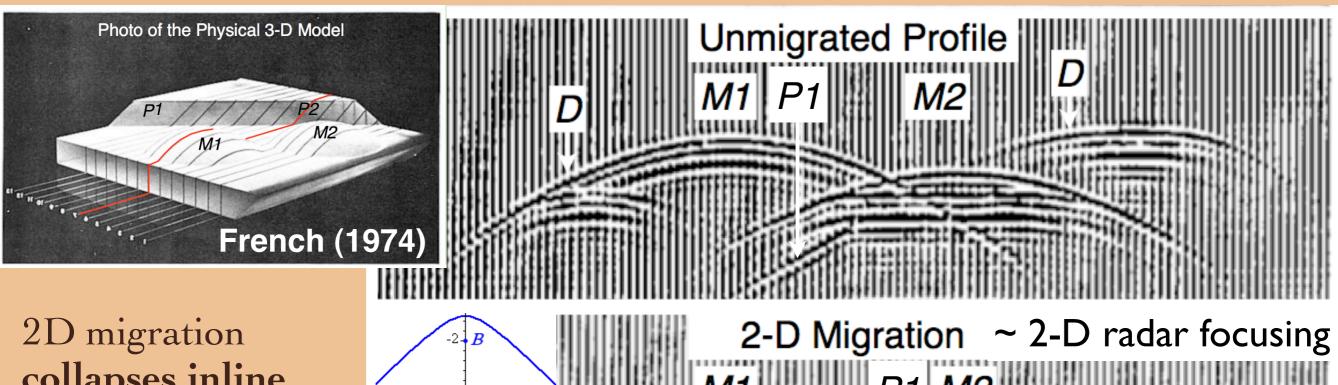
Diffractions (D) occur at edges of slope breaks.

Clutter from out-ofplane structures (M2, P1) interferes with nadir returns.

No returns from features oriented obliquely to the profile **(P2)**.

Photo of the Physical 3-D Model Profile 7 crosses one mound, skirts another, then obliquely crosses a slope. Mi French (1974) **Ultrasonic Profile 7**

Incomplete solution: Apply 2D migration*



2D migration collapses inline diffractions (D), but it mispositions out-of-plane returns (M2, P1).

e), ions P1 M2 (1).

*Migration works in a moving window of data, summing values along an operator (in 2D, a hyperbola in a constant-velocity medium), essentially treating each point in the dataset as a potential diffractor.

Usage note: *Migration* in processing is unrelated to trough *migration*!

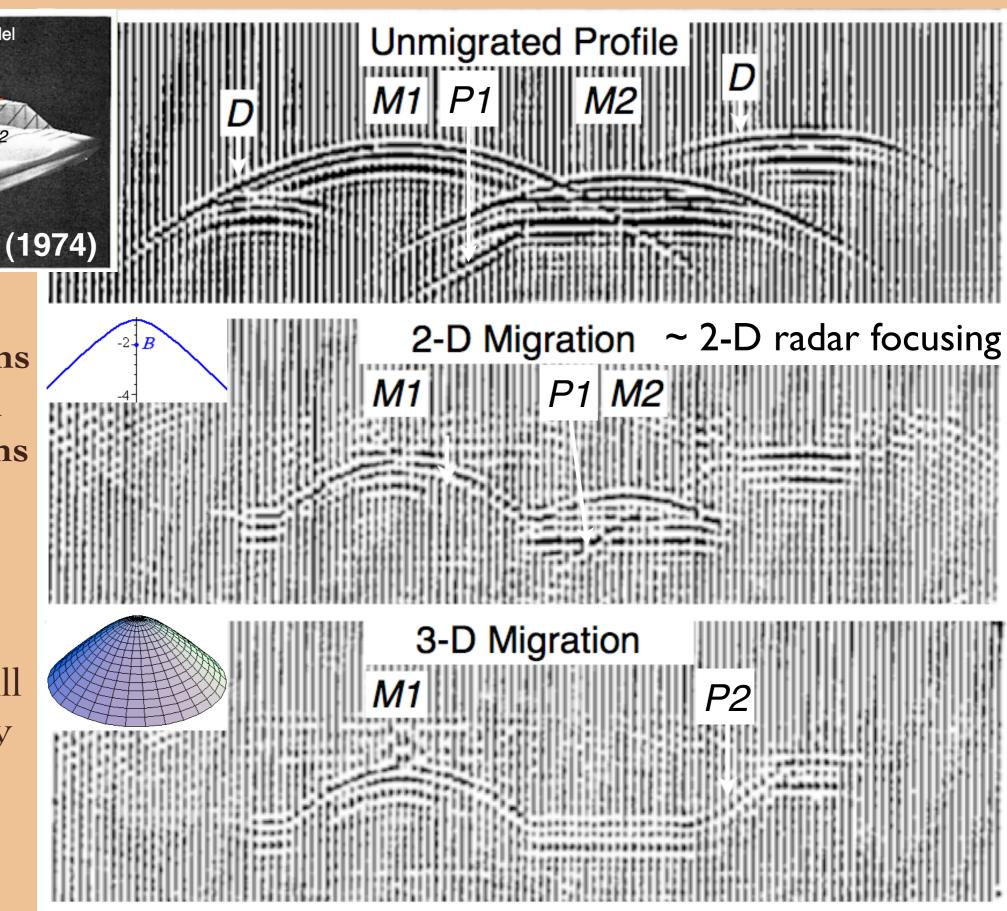
Full solution: Apply 3D migration to all data

Photo of the Physical 3-D Model

French (1974)

3-D migration will
collapse diffractions
(D) and reposition
out-of-plane returns
(M2, P1) to their
source locations.

Energy from adjacent profiles will be restored, thereby **imaging features oriented obliquely to the profile (P2)**.

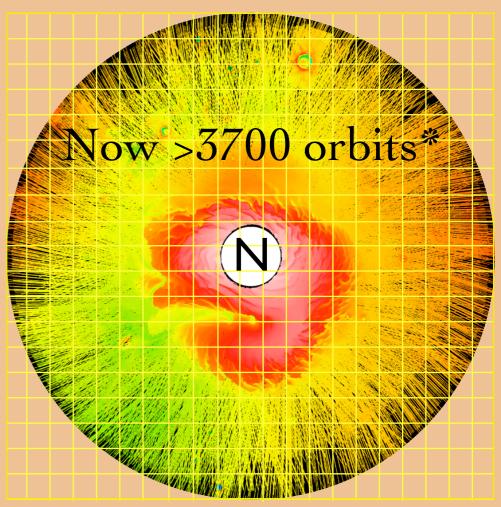


Creating a 3D volume from the 2D SHARAD data

- Apply 2D Doppler processing & autofocusing to correct ionospheric phase distortions.
- Apply demigration to reverse the 2D focusing.
- Adust the demigrated radargrams
 to a common datum* and bin them into a rectilinear grid.
- * Datuming step is complicated by alongtrack variable delays in the ionosphere.

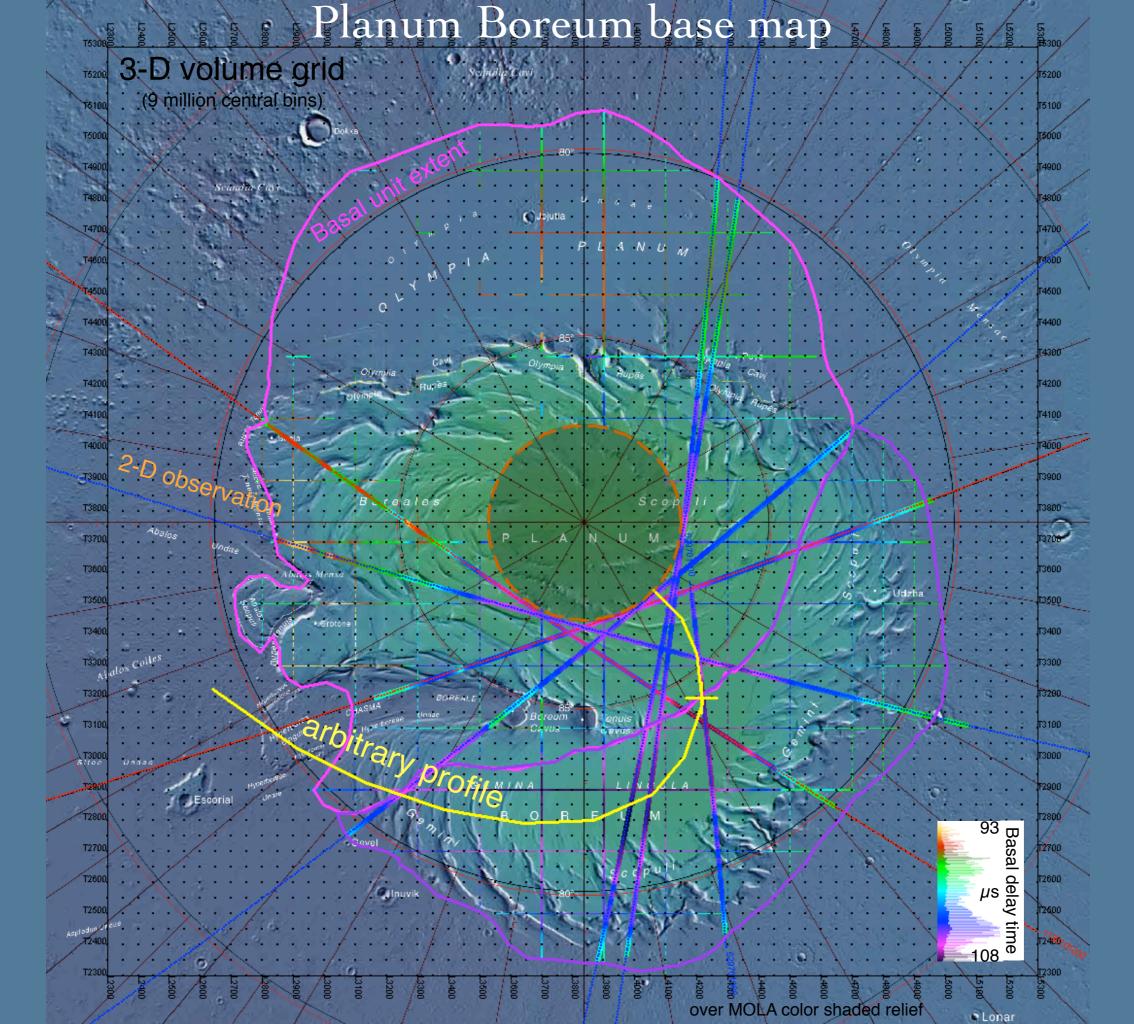
See Campbell et al. (GRSL 2014)

Coverage 70–87°N (in MOLA elevation)



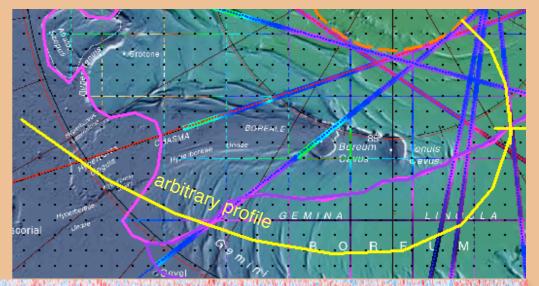
* 3-D grid covers 6073 × 5489 km,
133 million bins of 500 × 500 m,
including 1579 observations.

Larger than any terrestrial 3-D survey!



3D allows new perspectives

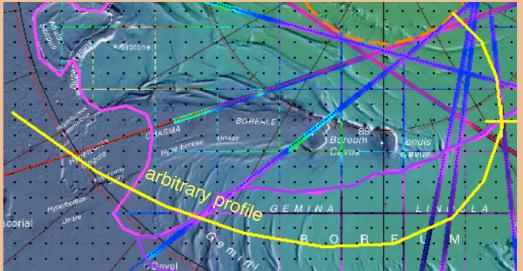
• Transect along Gemina Lingula apex provides a view of the data orthogonal to most of the troughs.

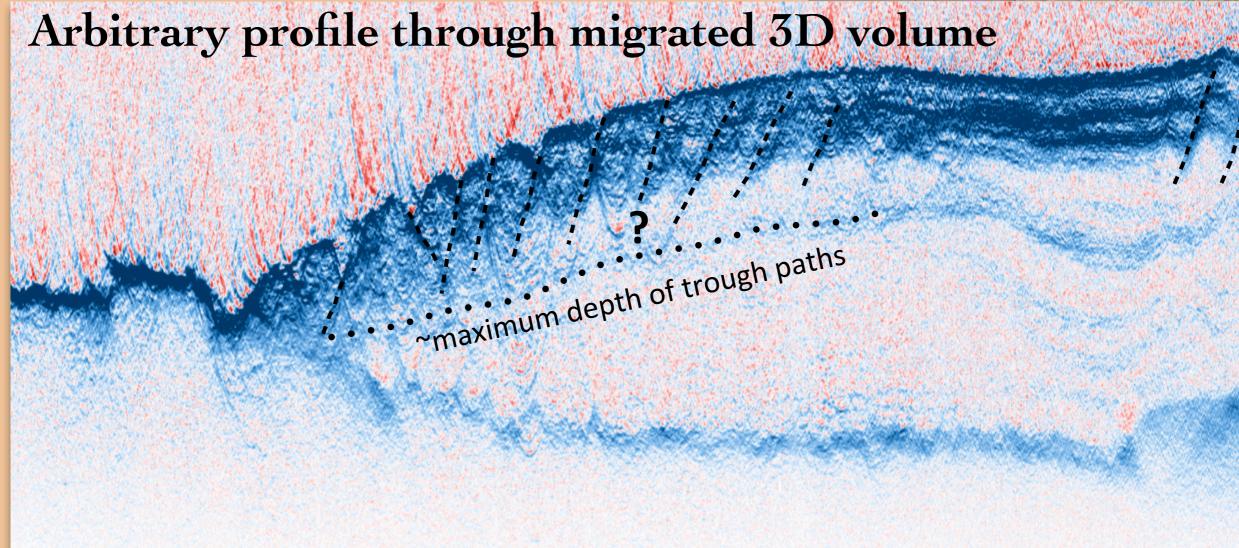


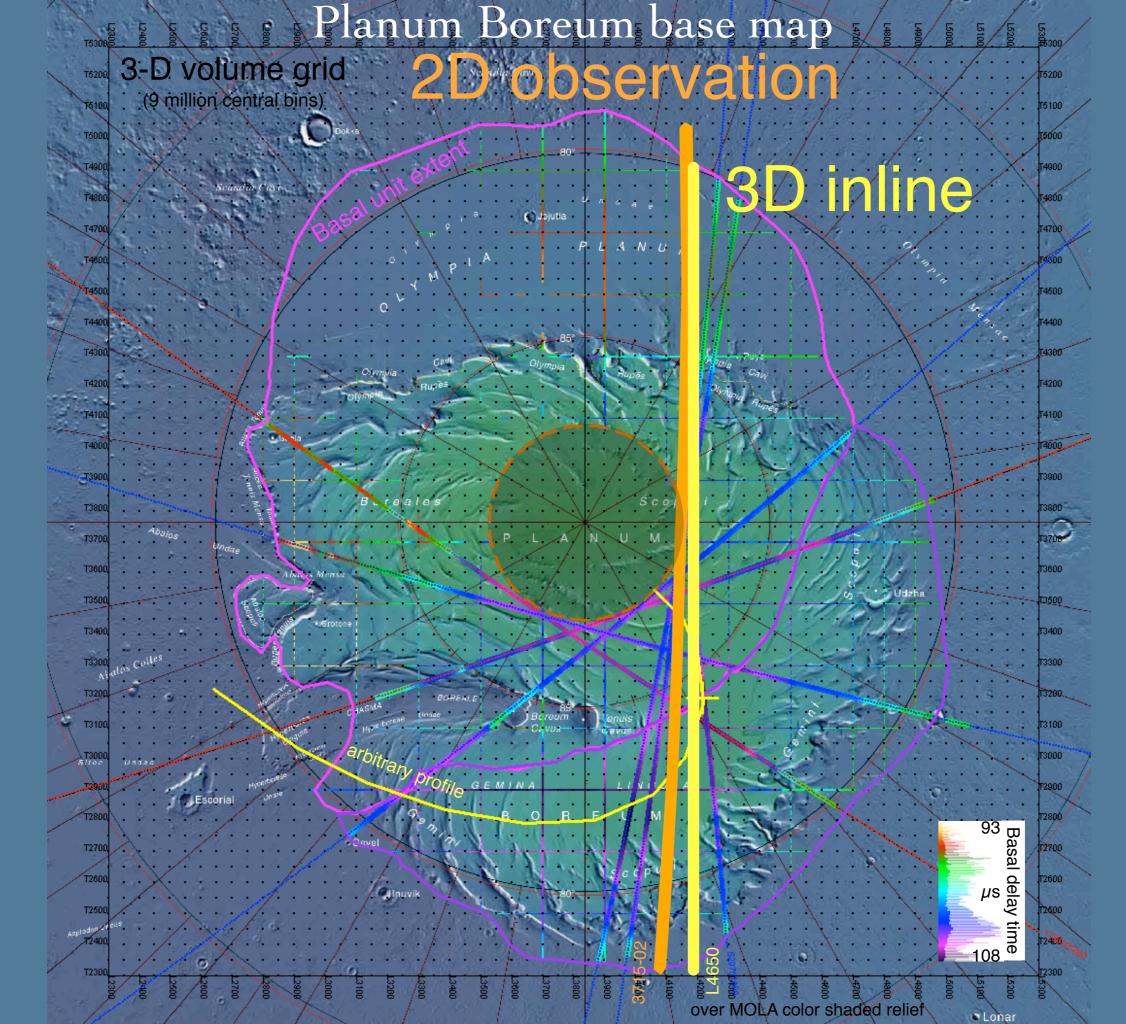
Arbitrary profile through migrated 3D volume

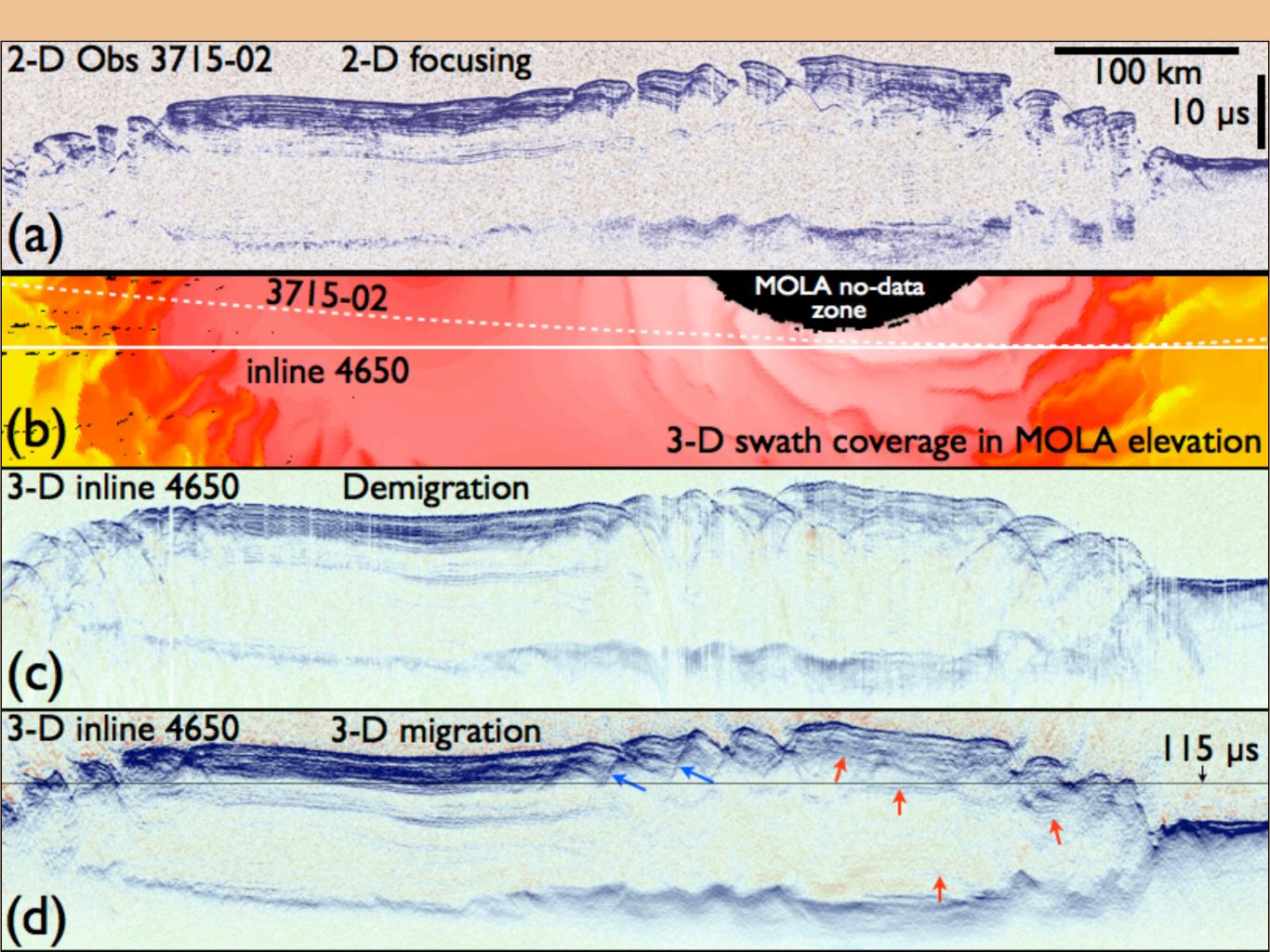
3D allows new perspectives

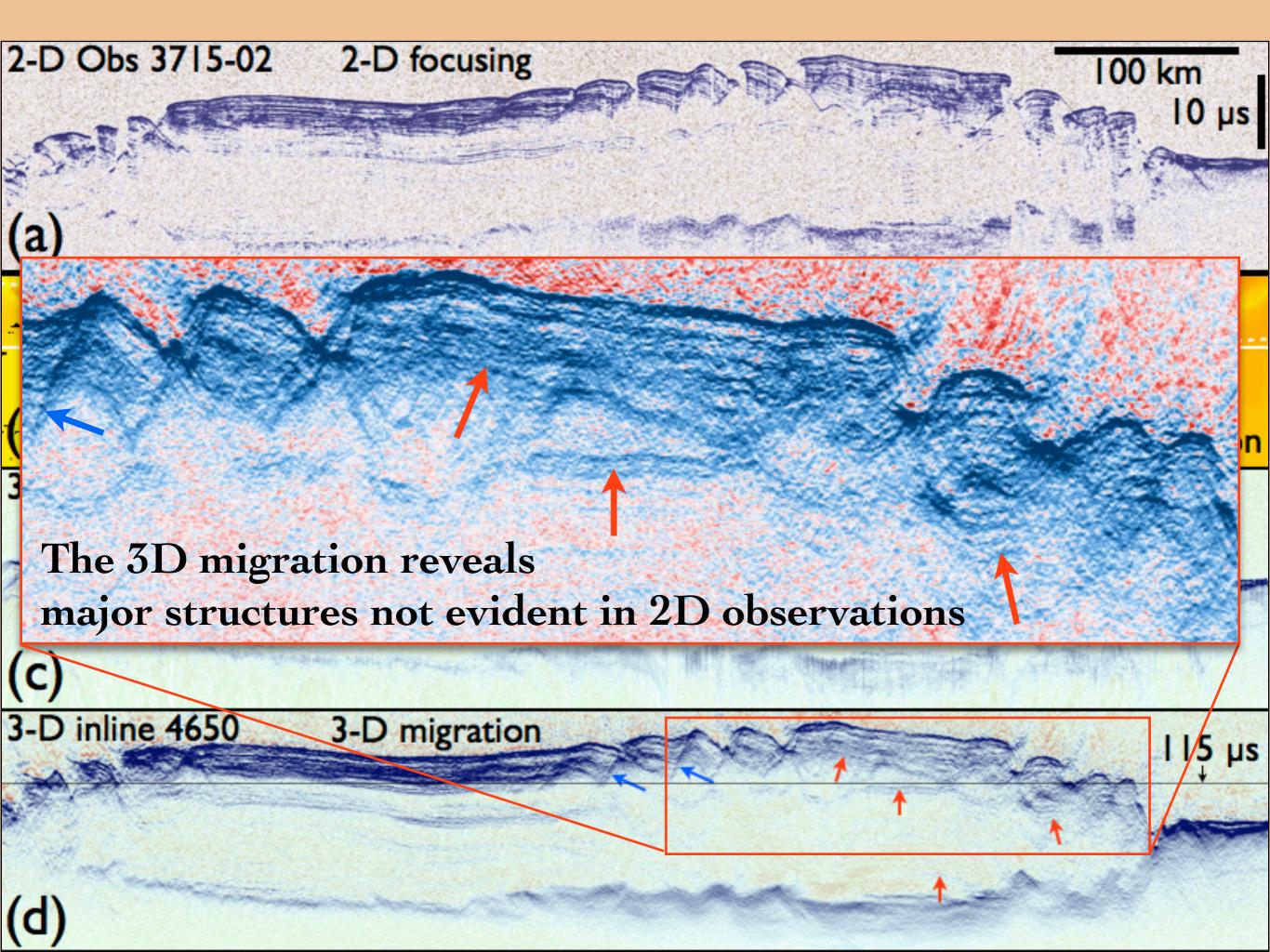
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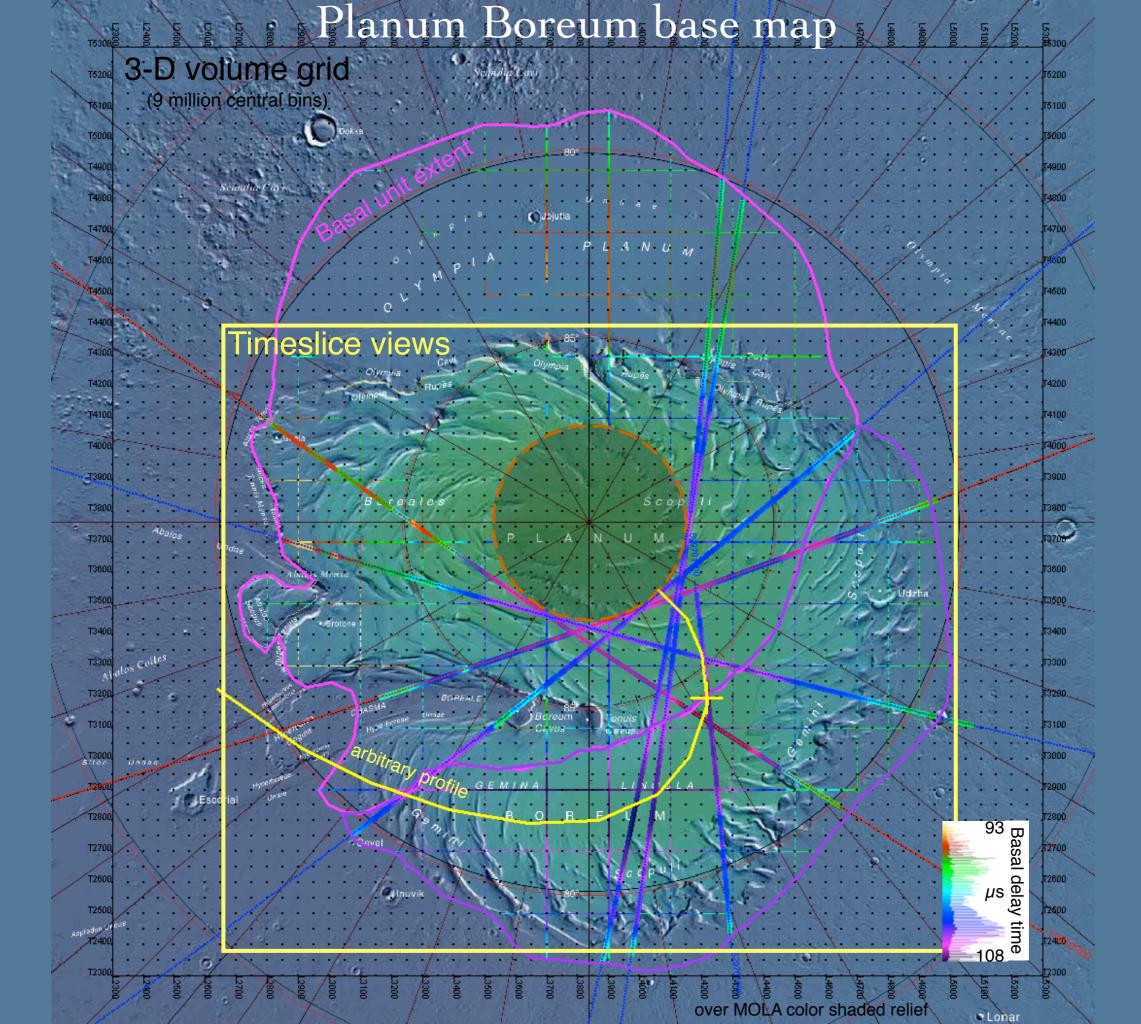


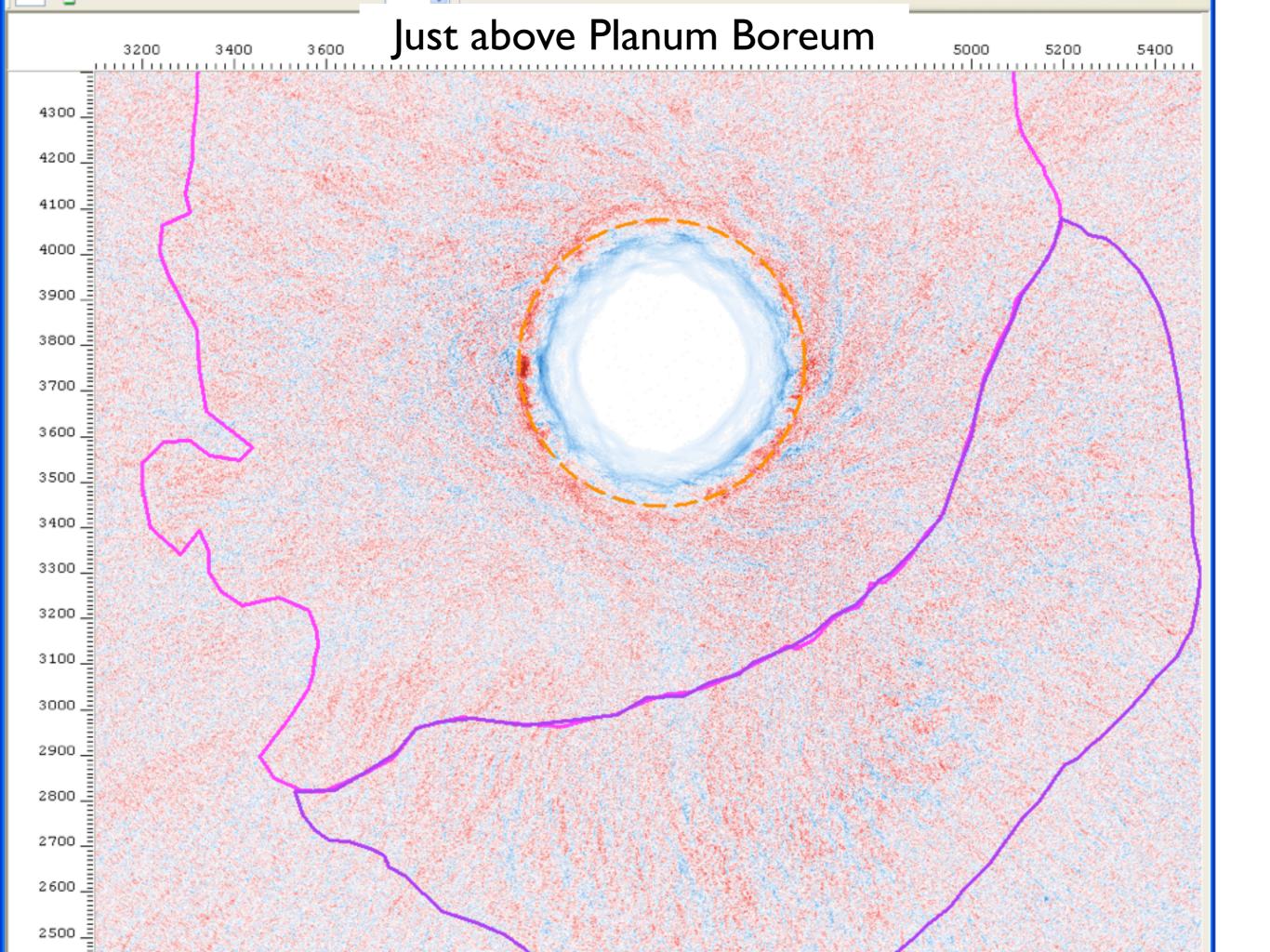


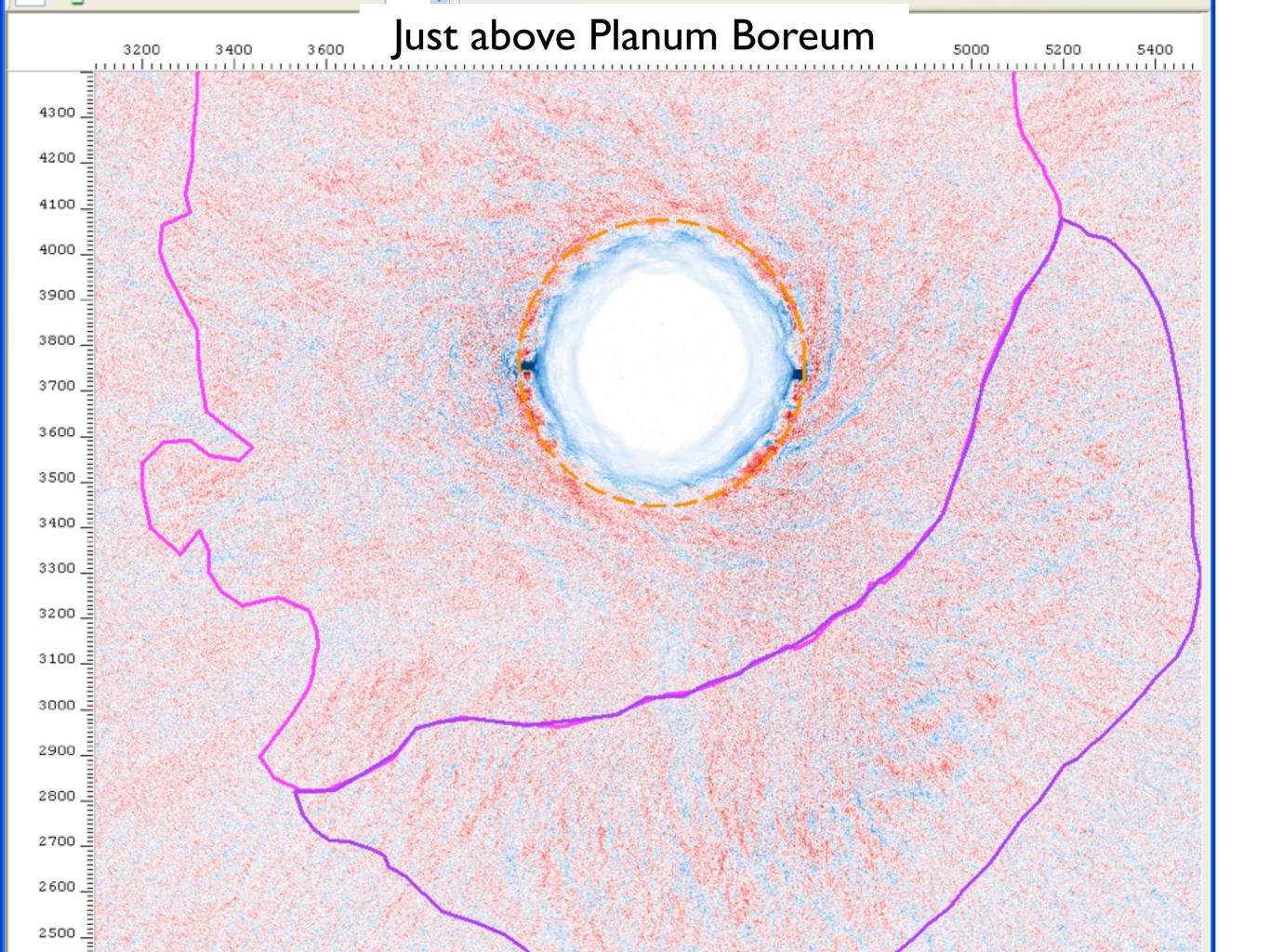


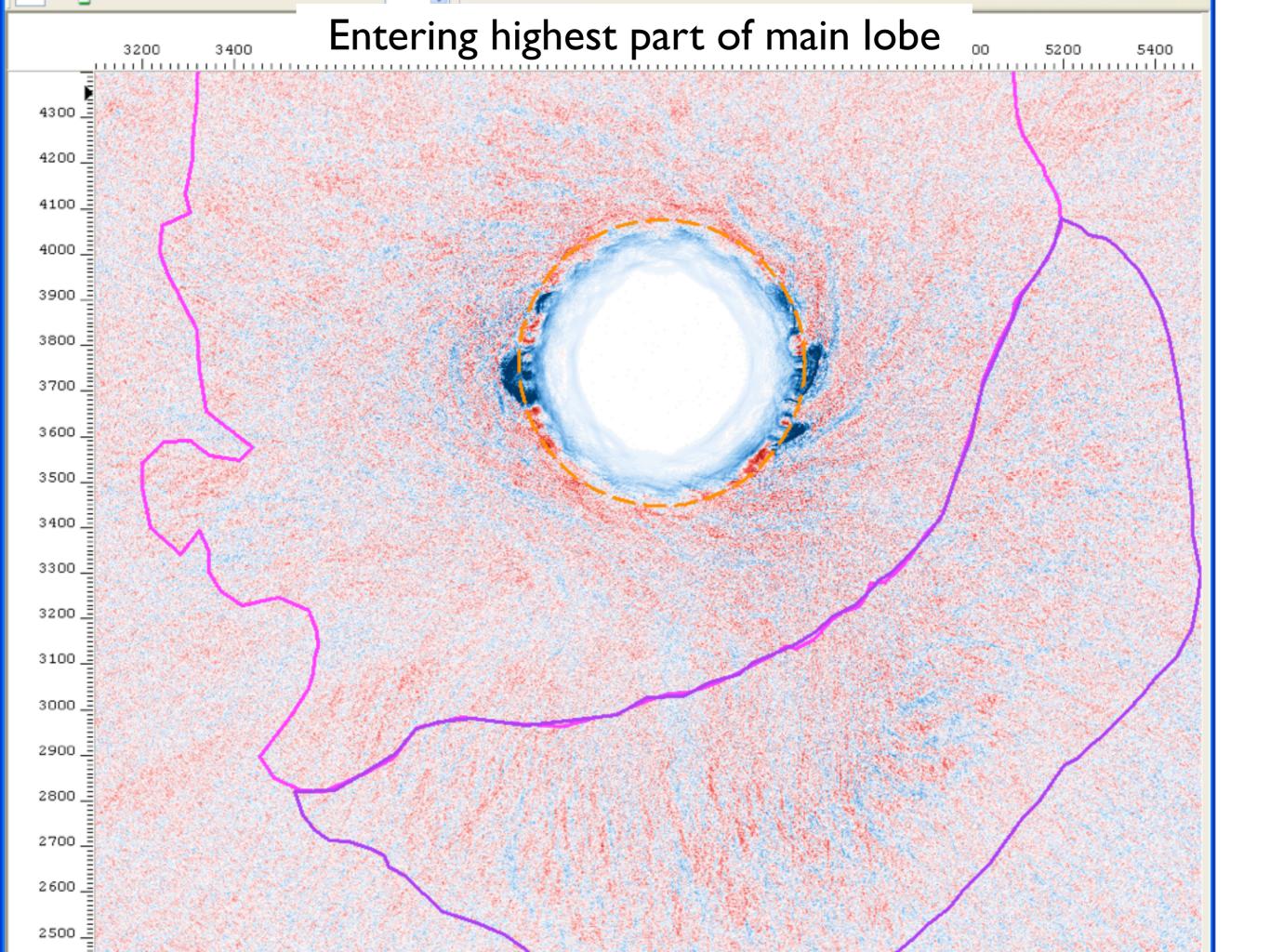


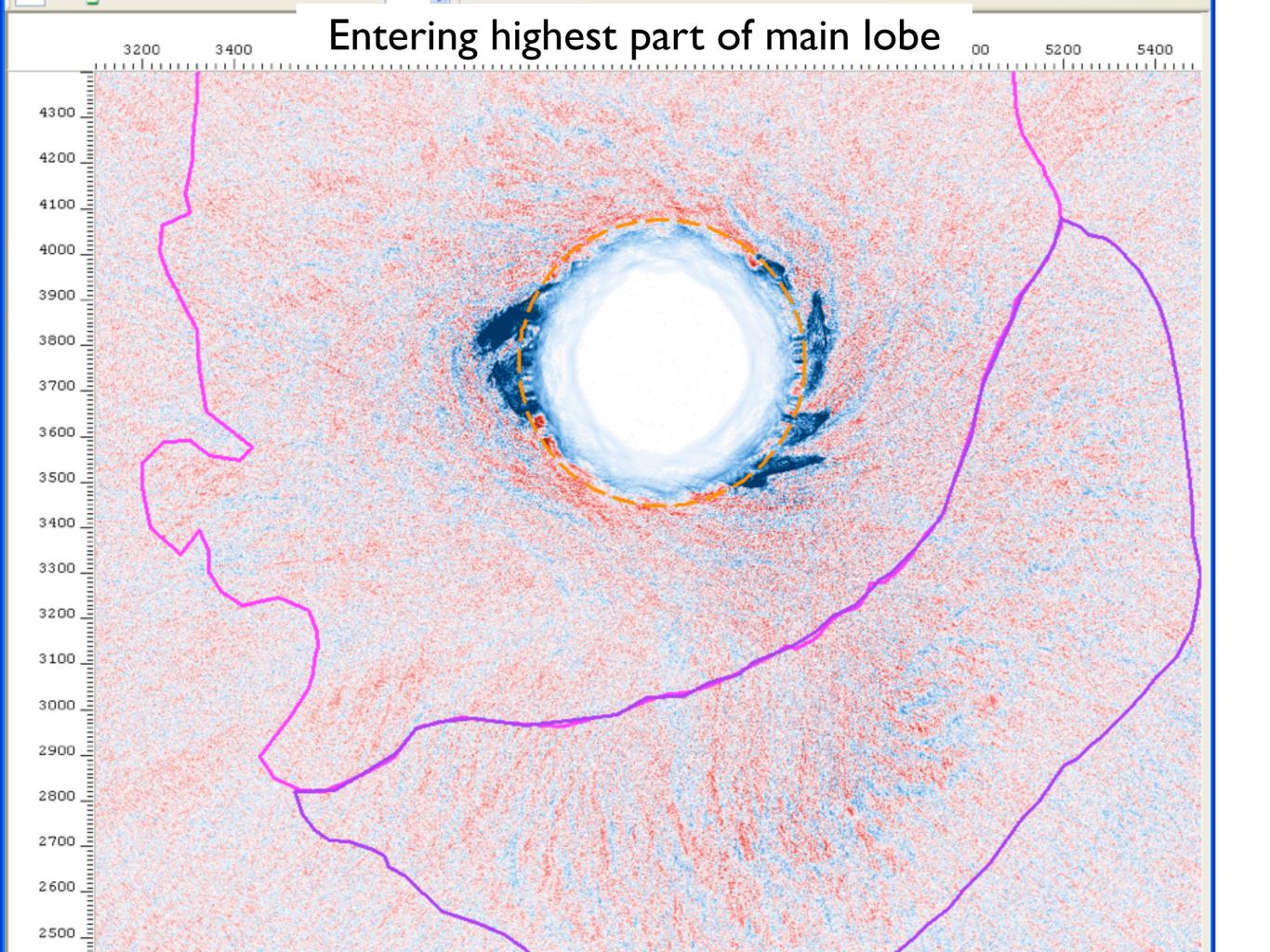


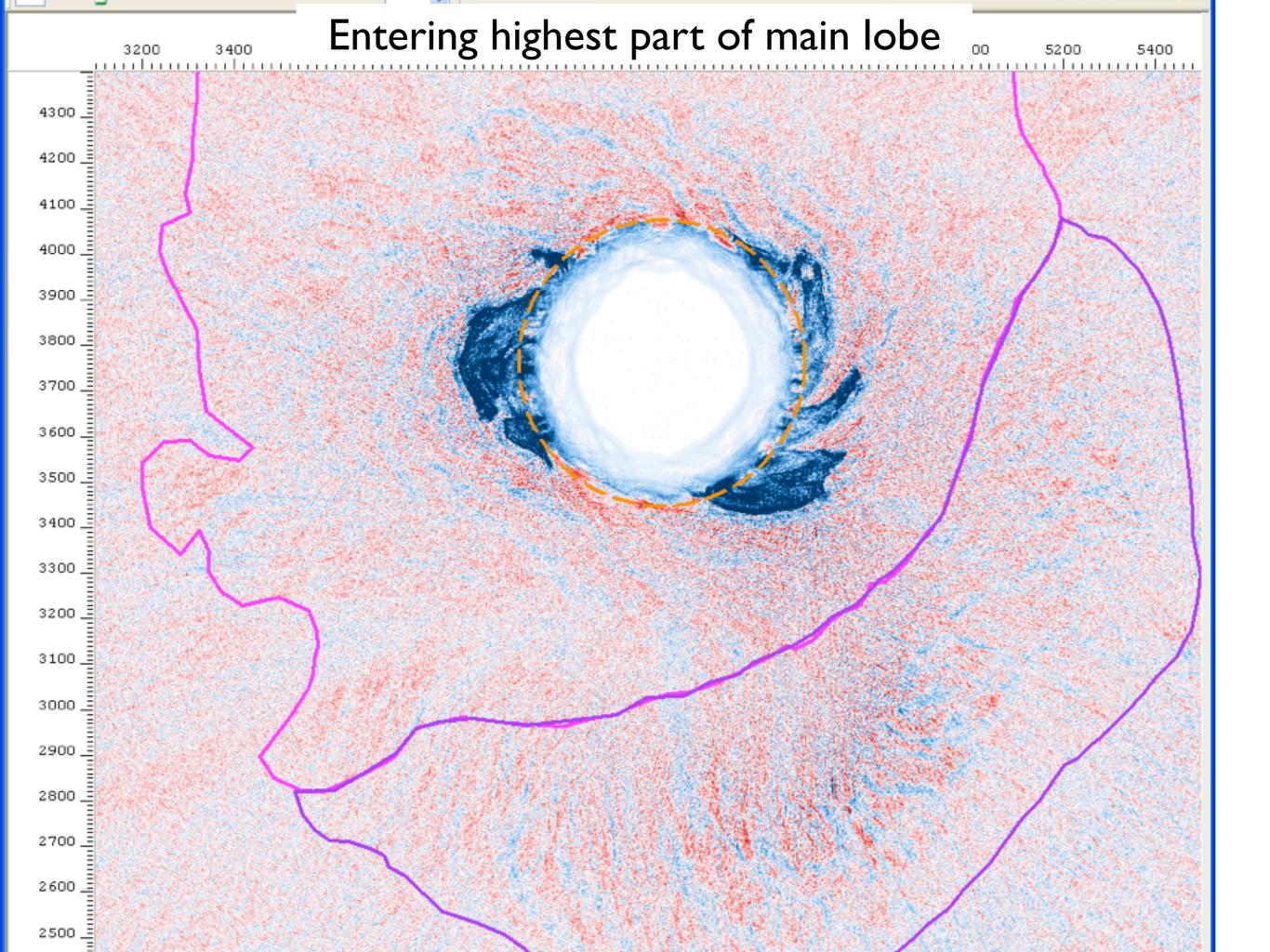


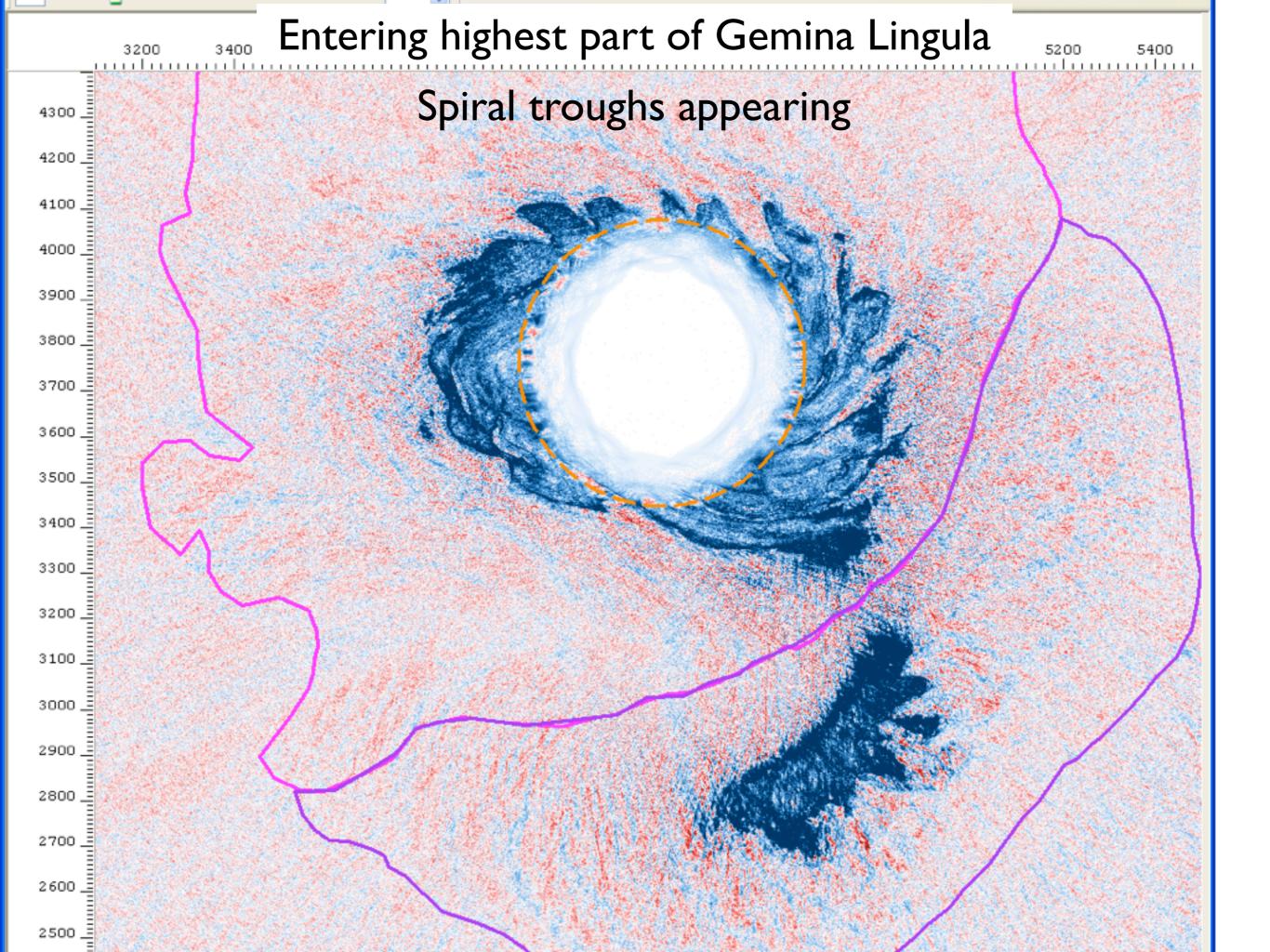


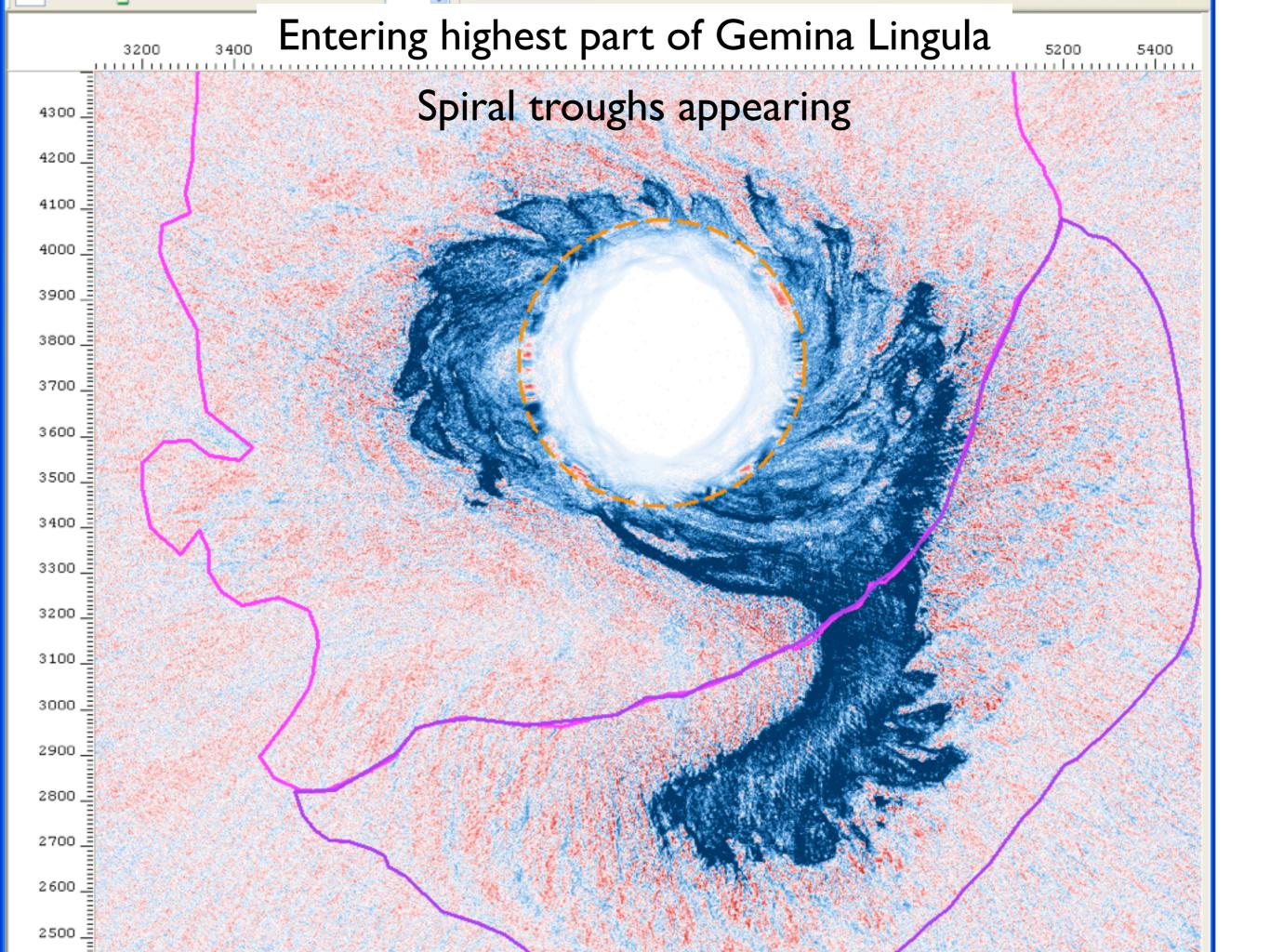


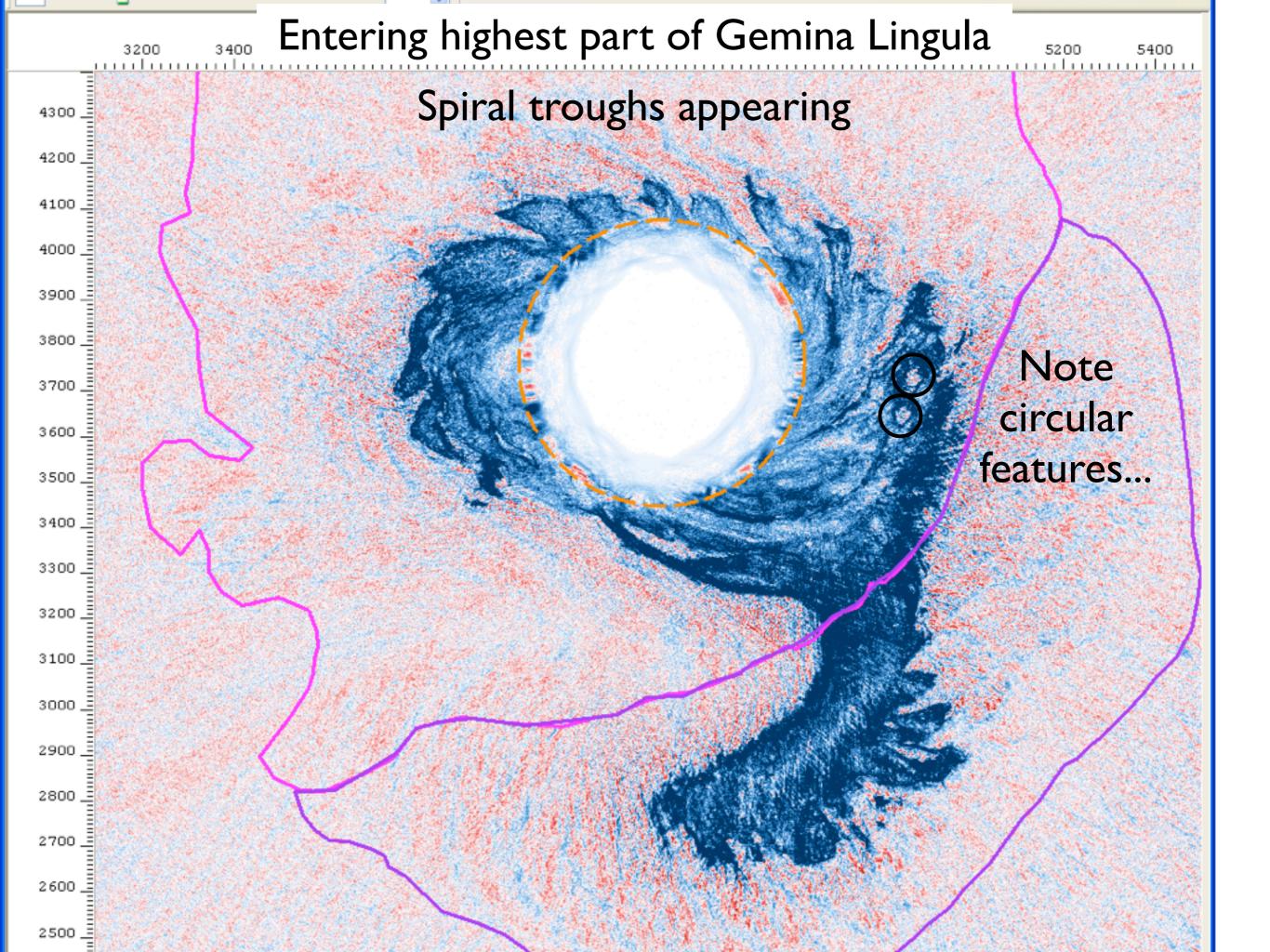


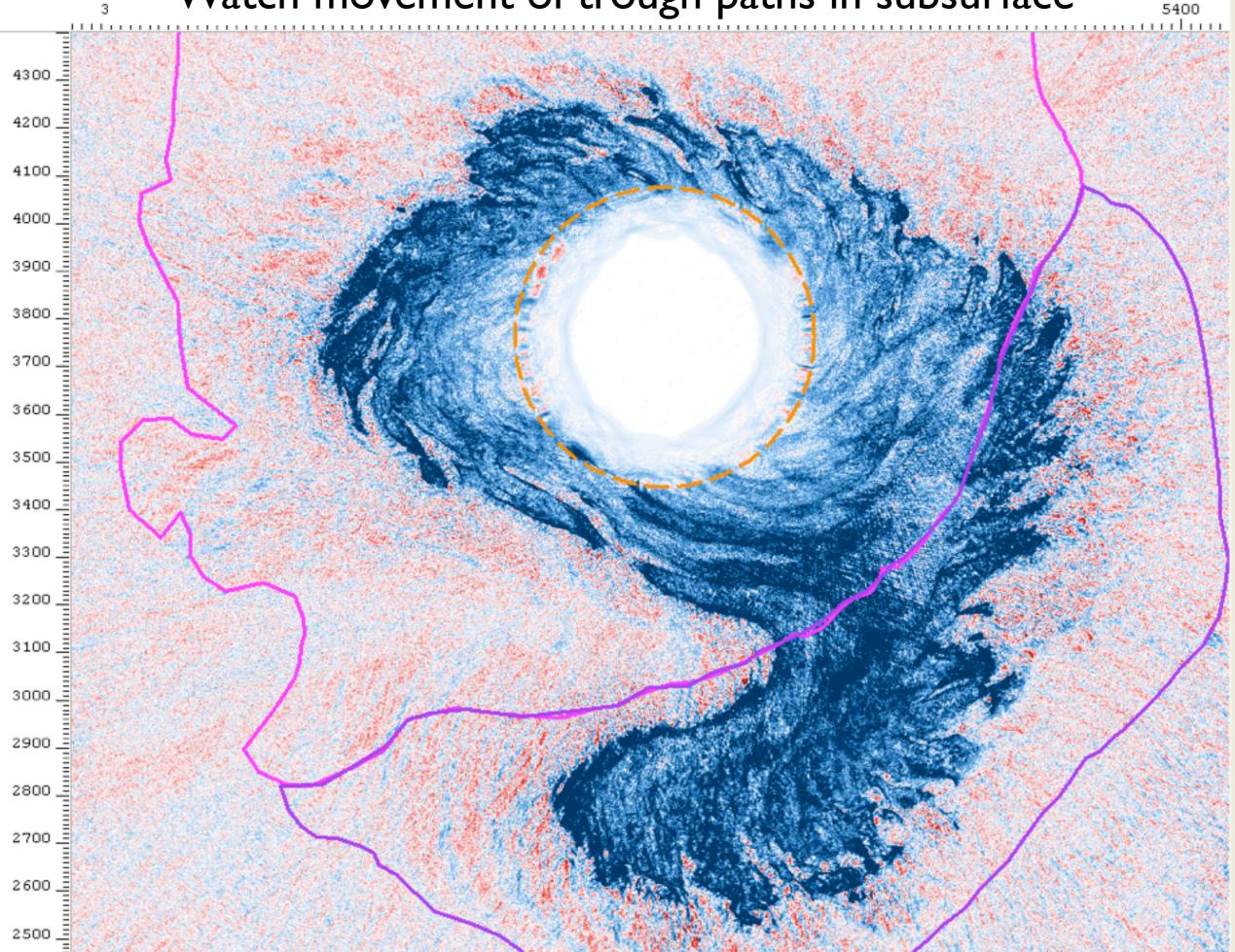


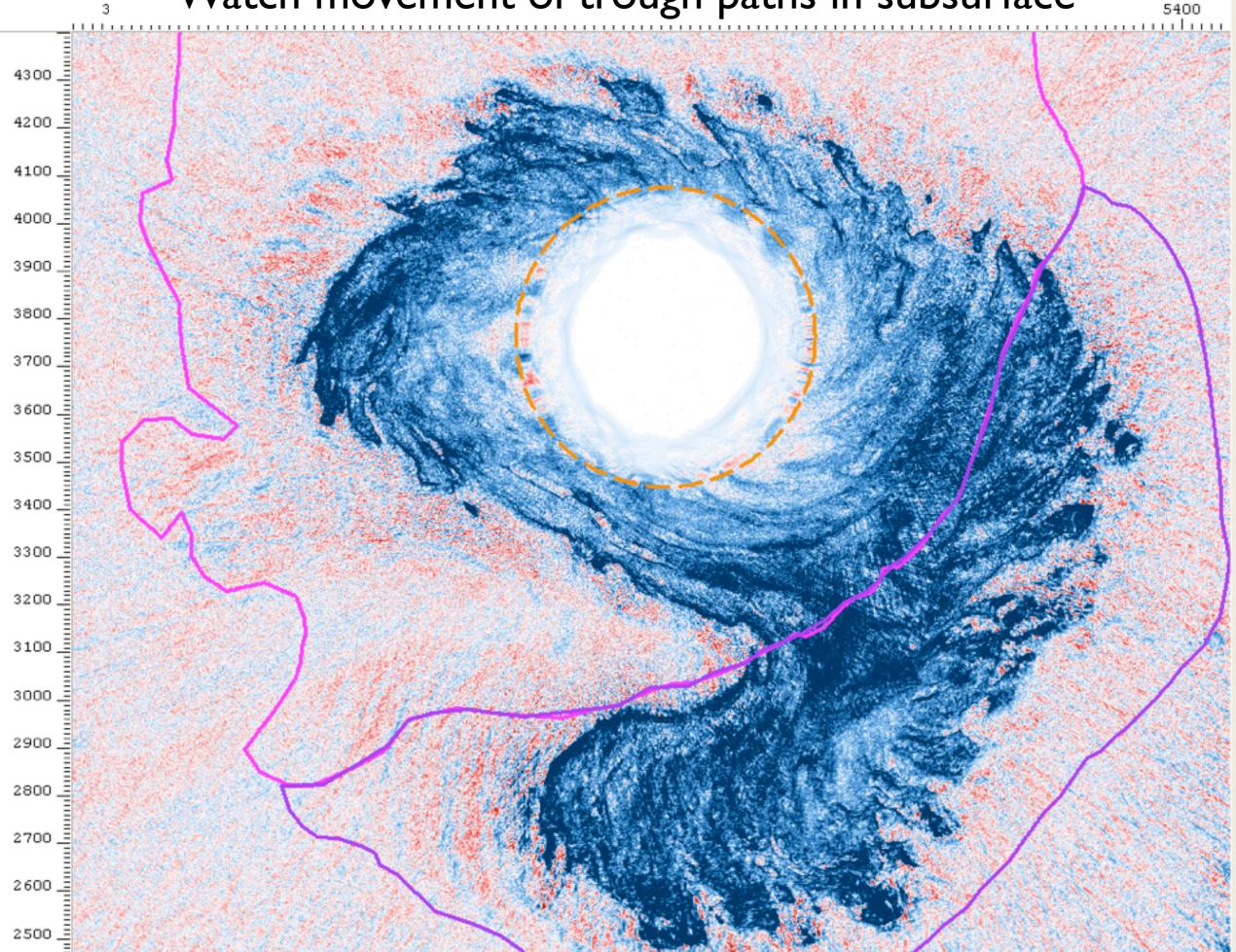




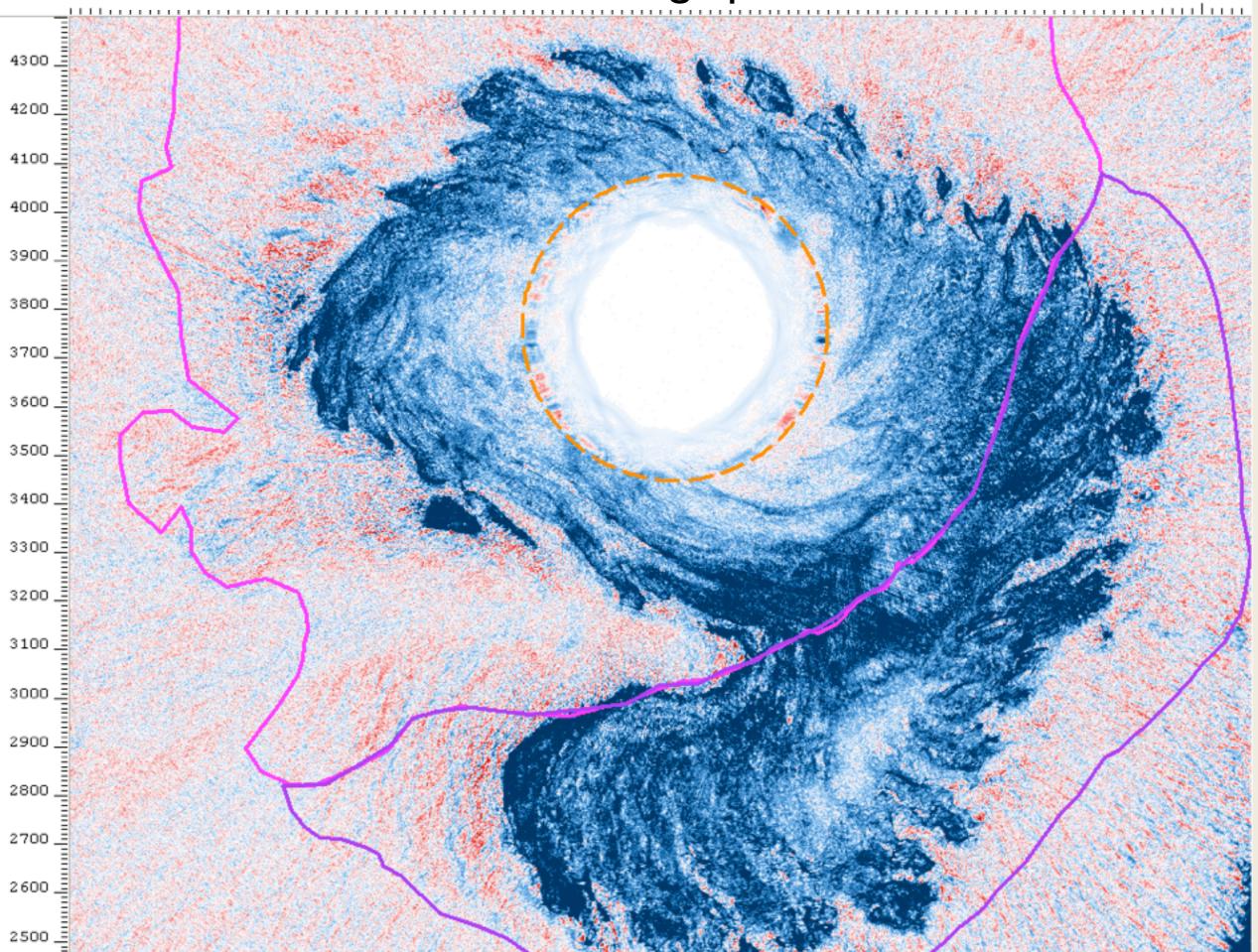




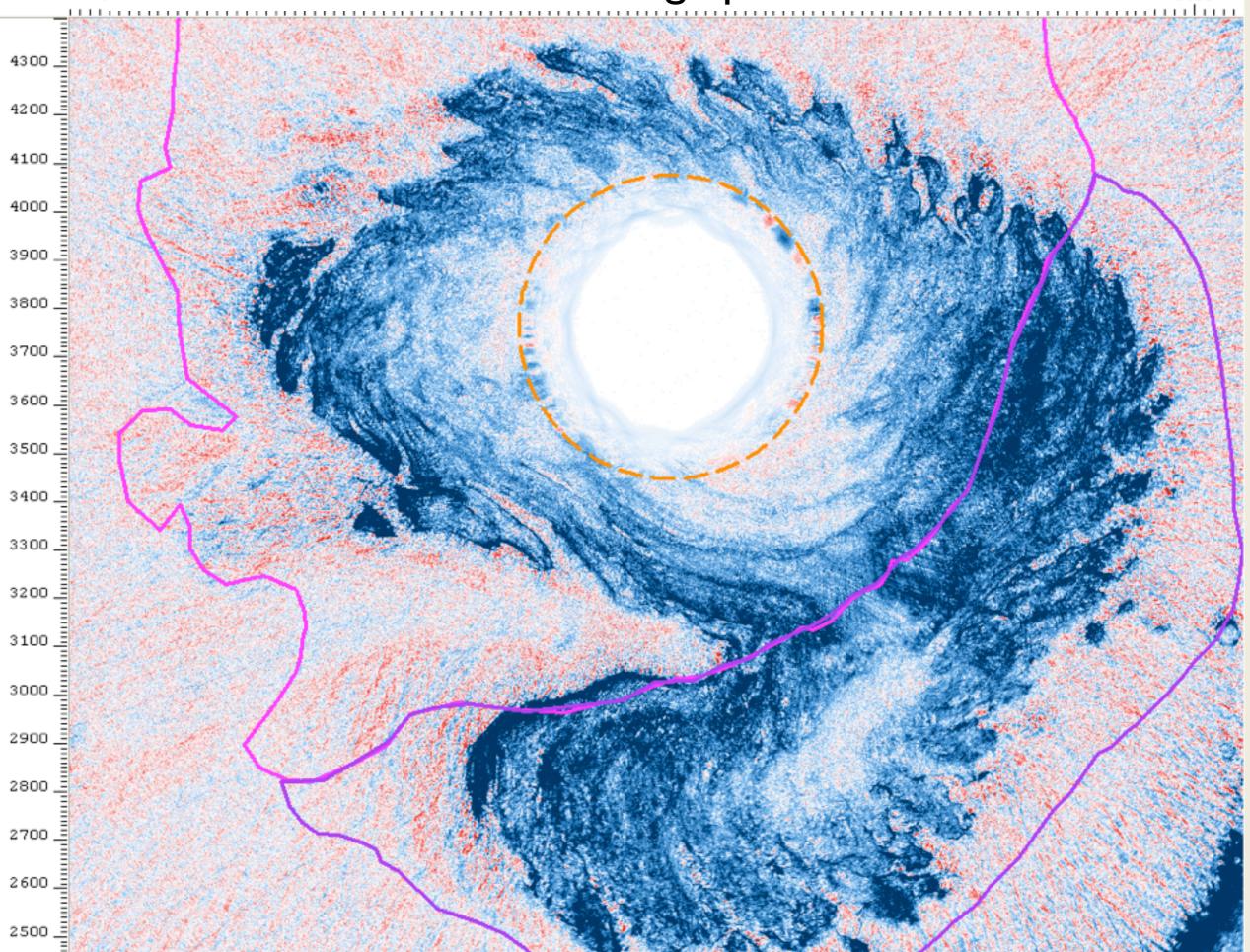


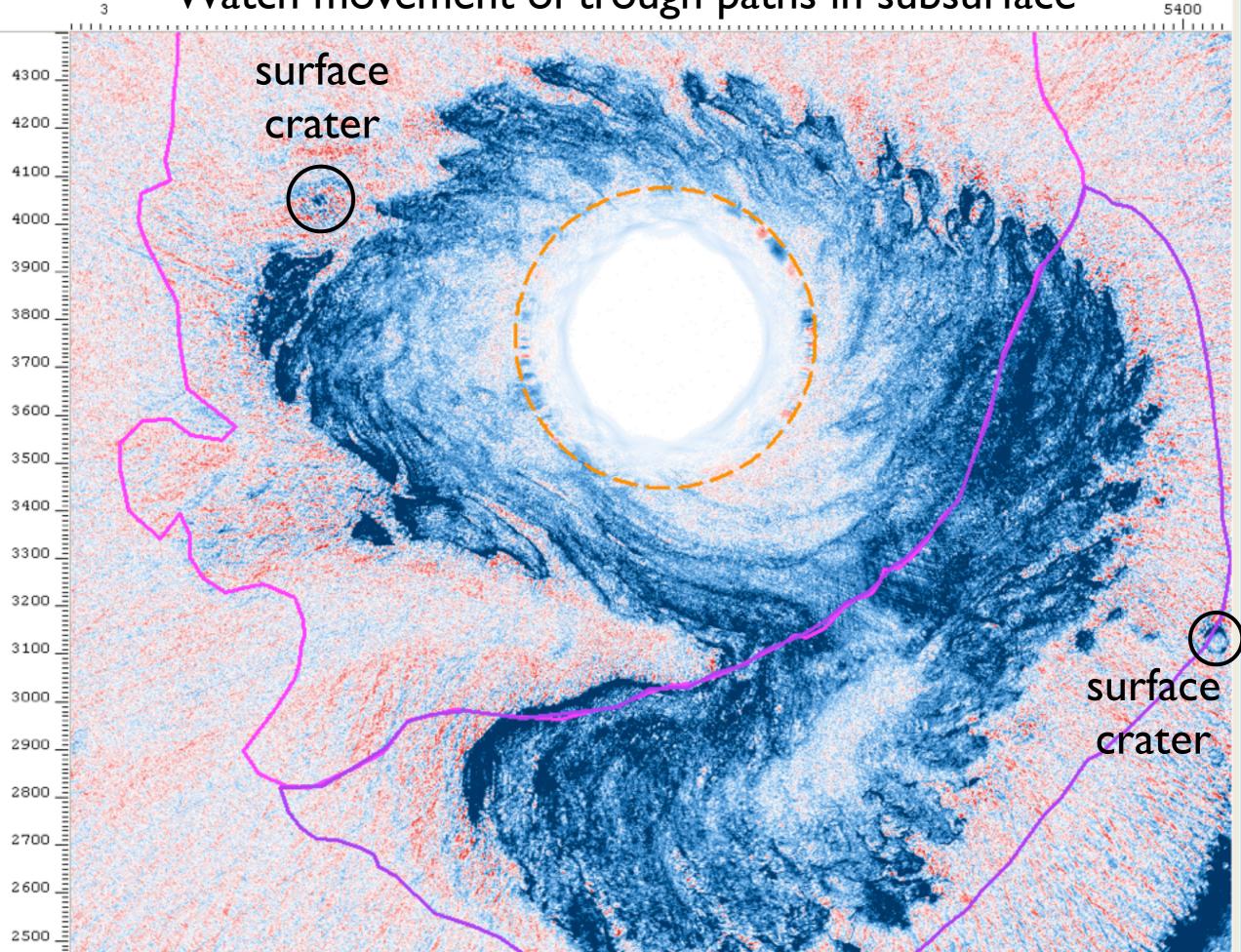


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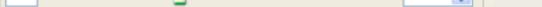


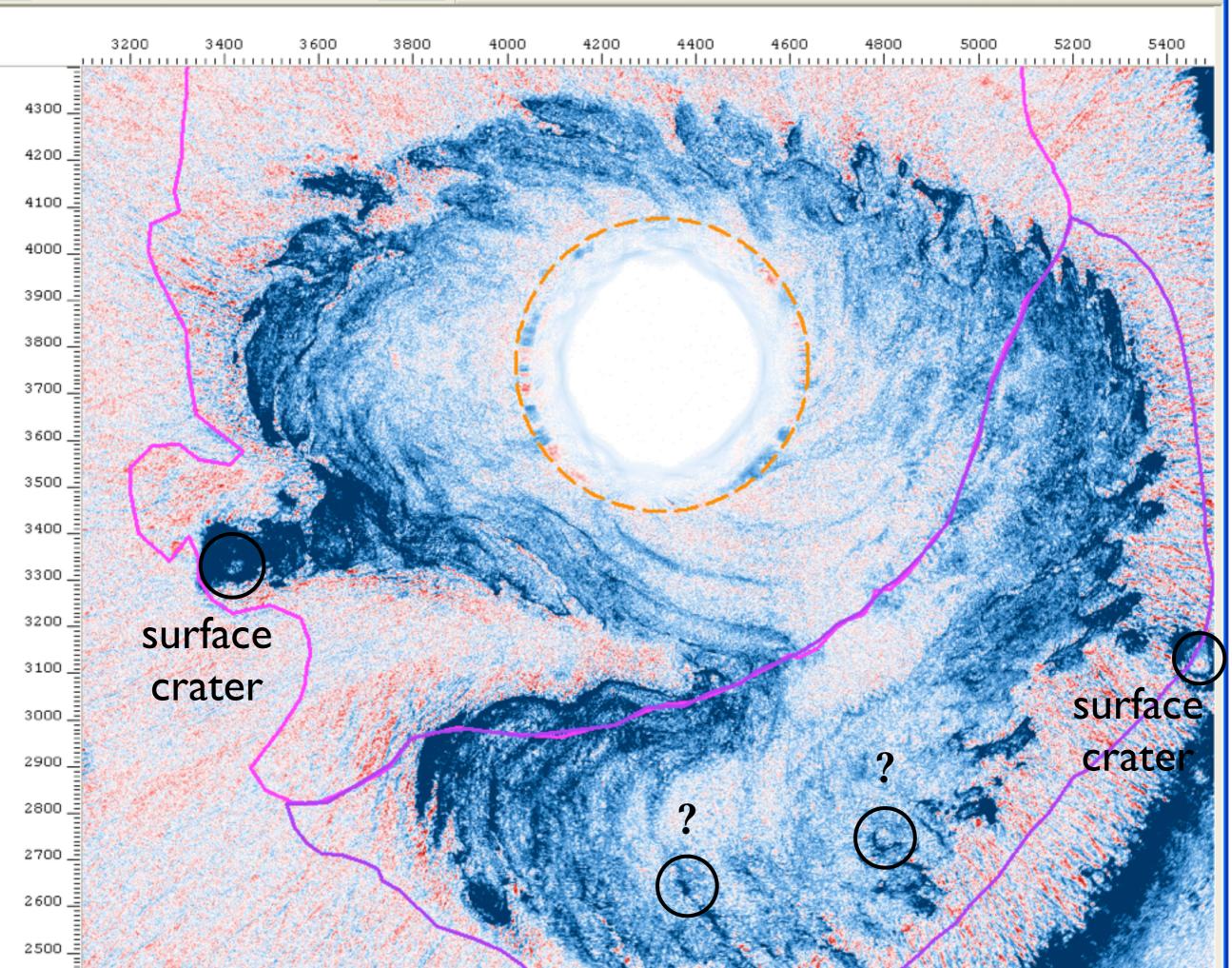


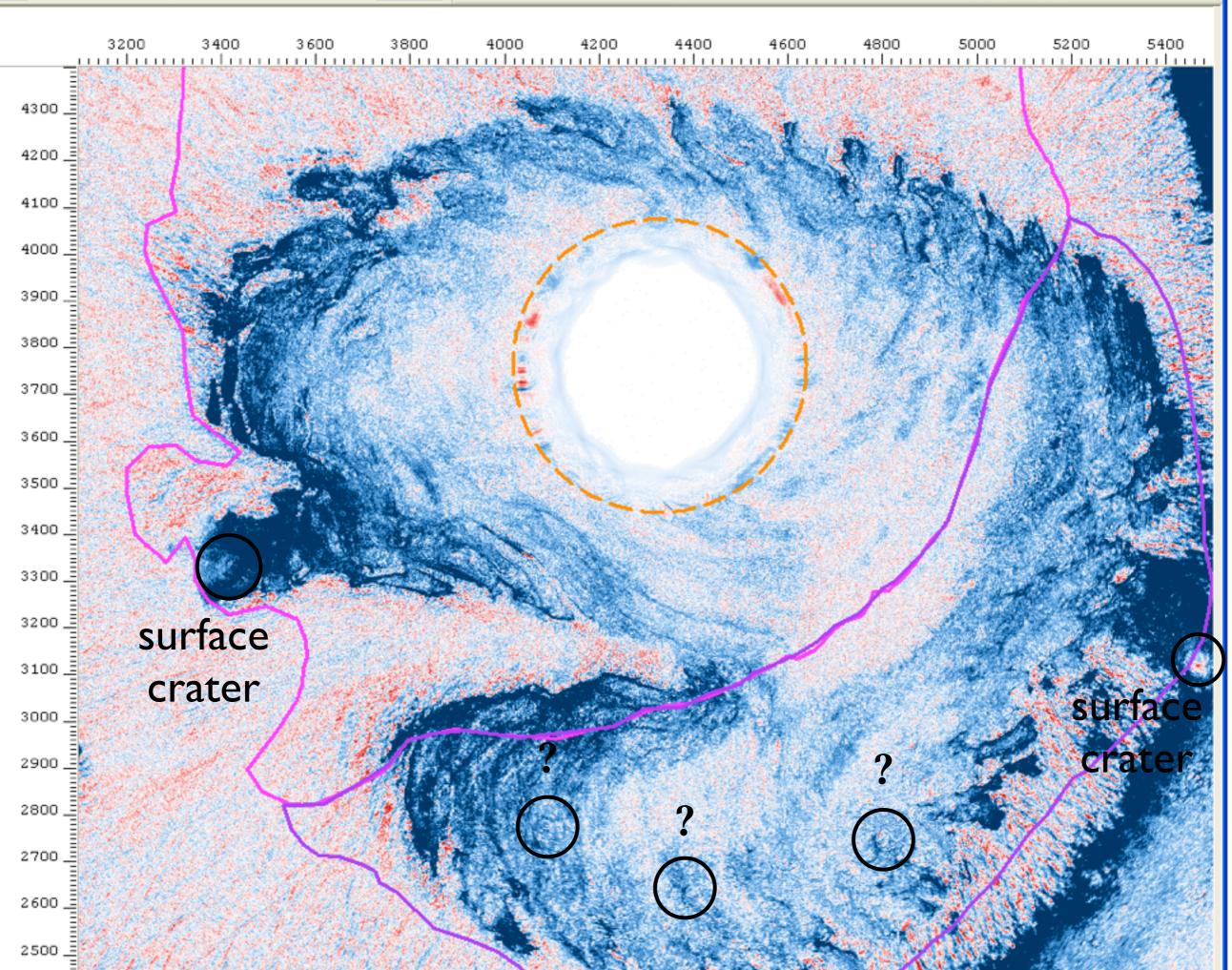
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2500 🚊

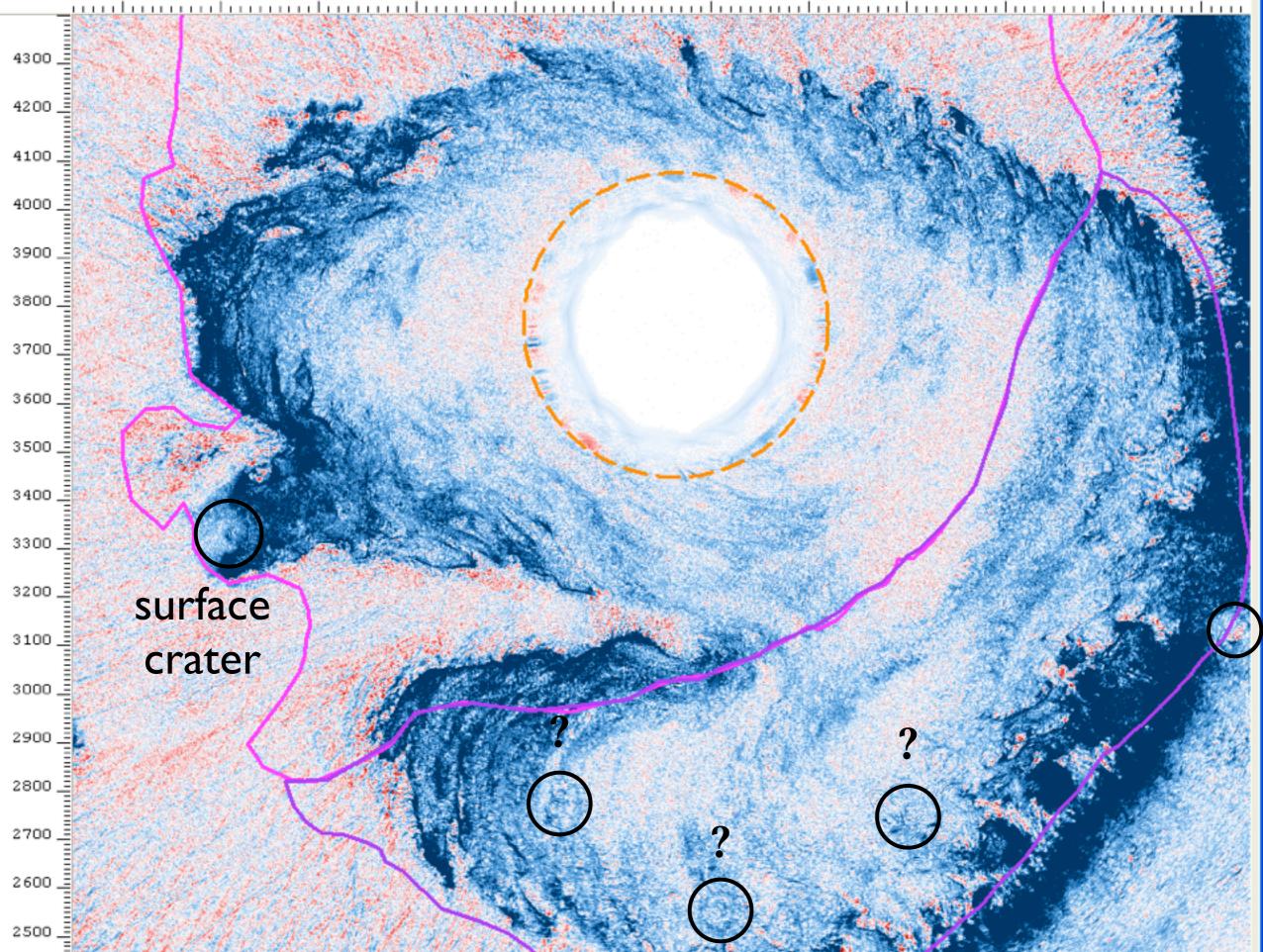
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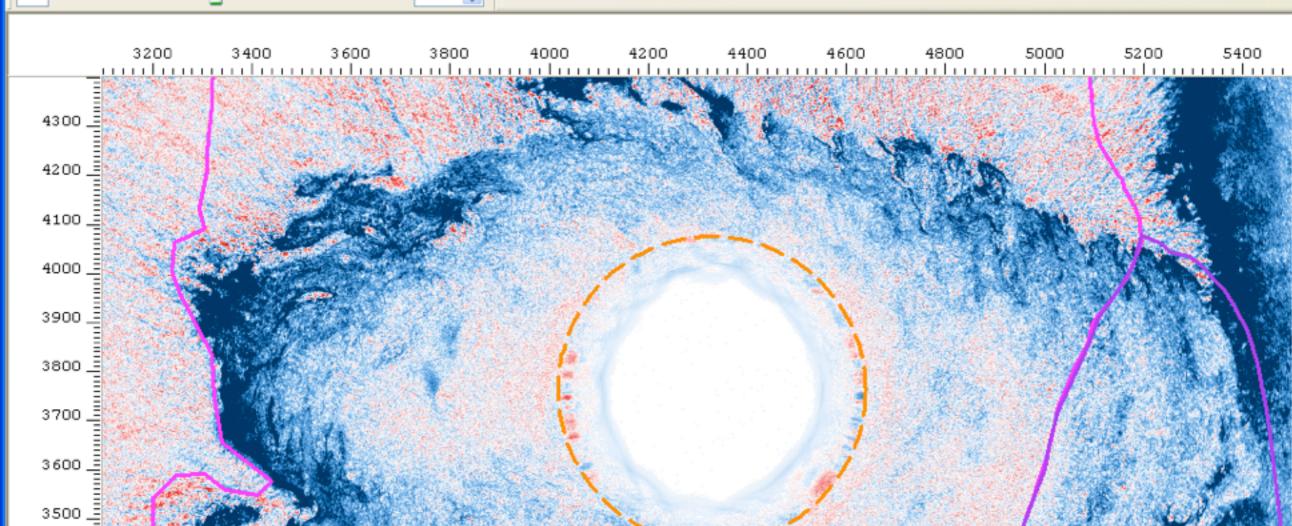






3200 3400 3600 3800 4000 4200 4400 4600 4800 5000 5200 5400





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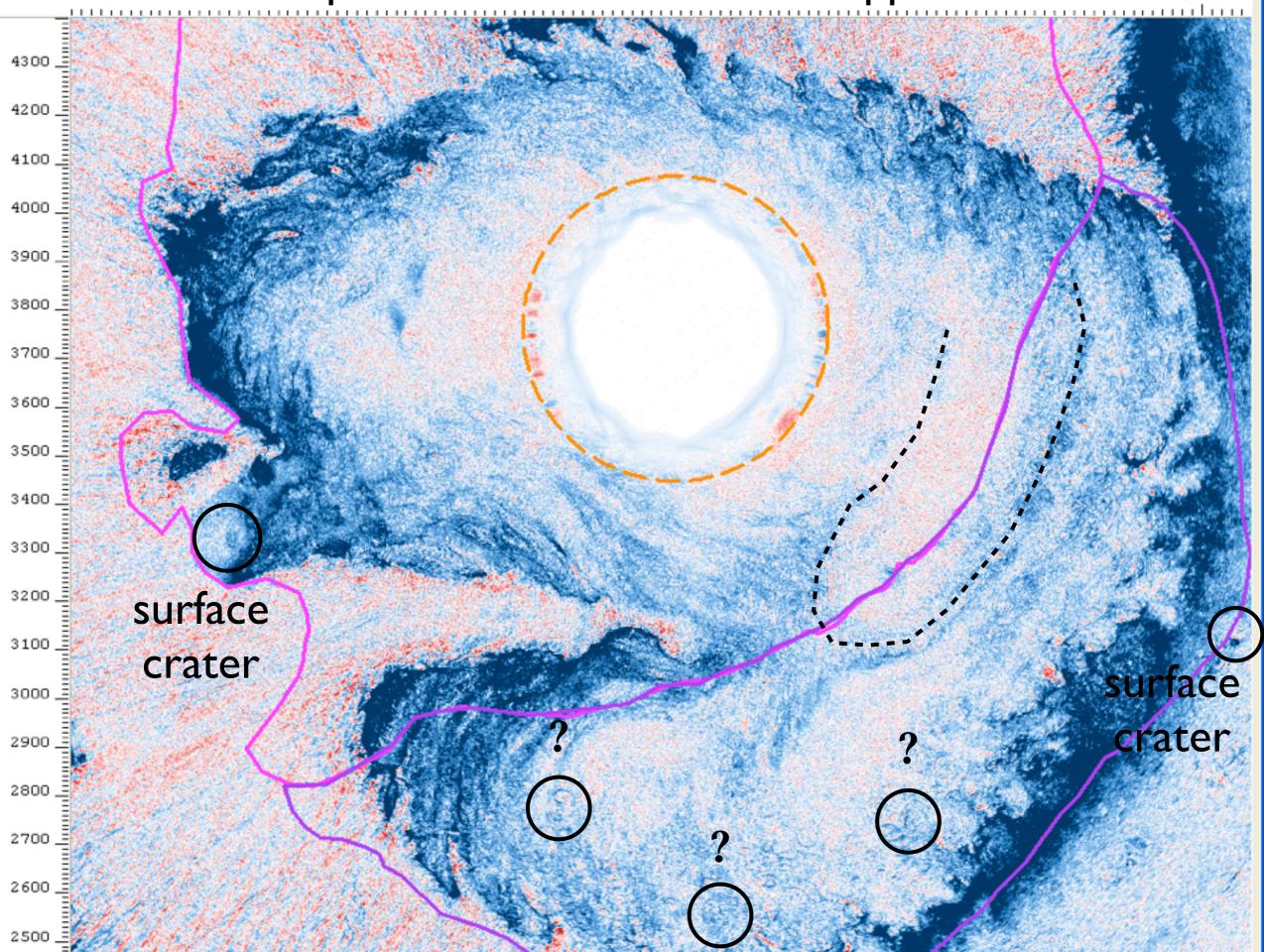
2500

surface

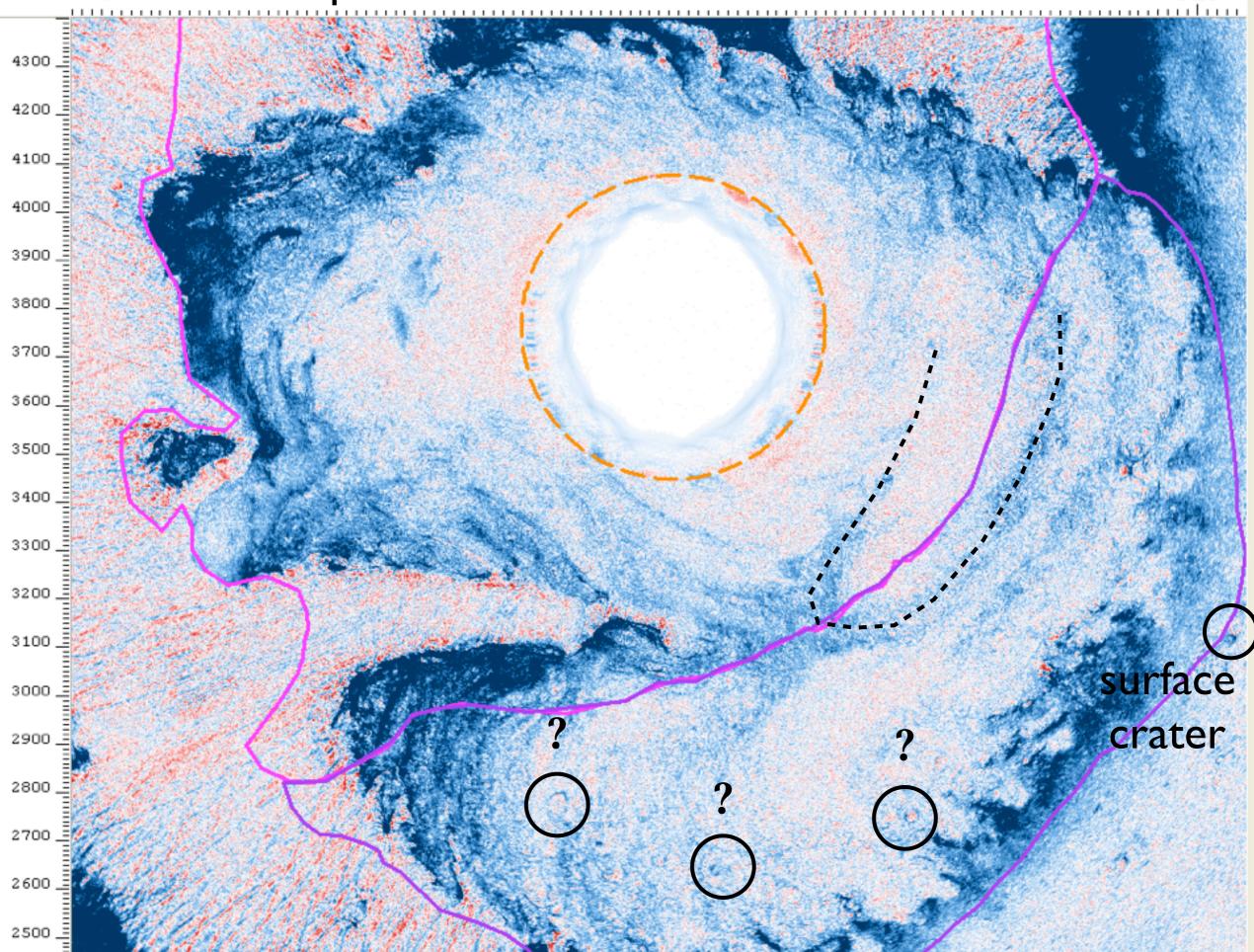
crater



Expression of buried chasma appears

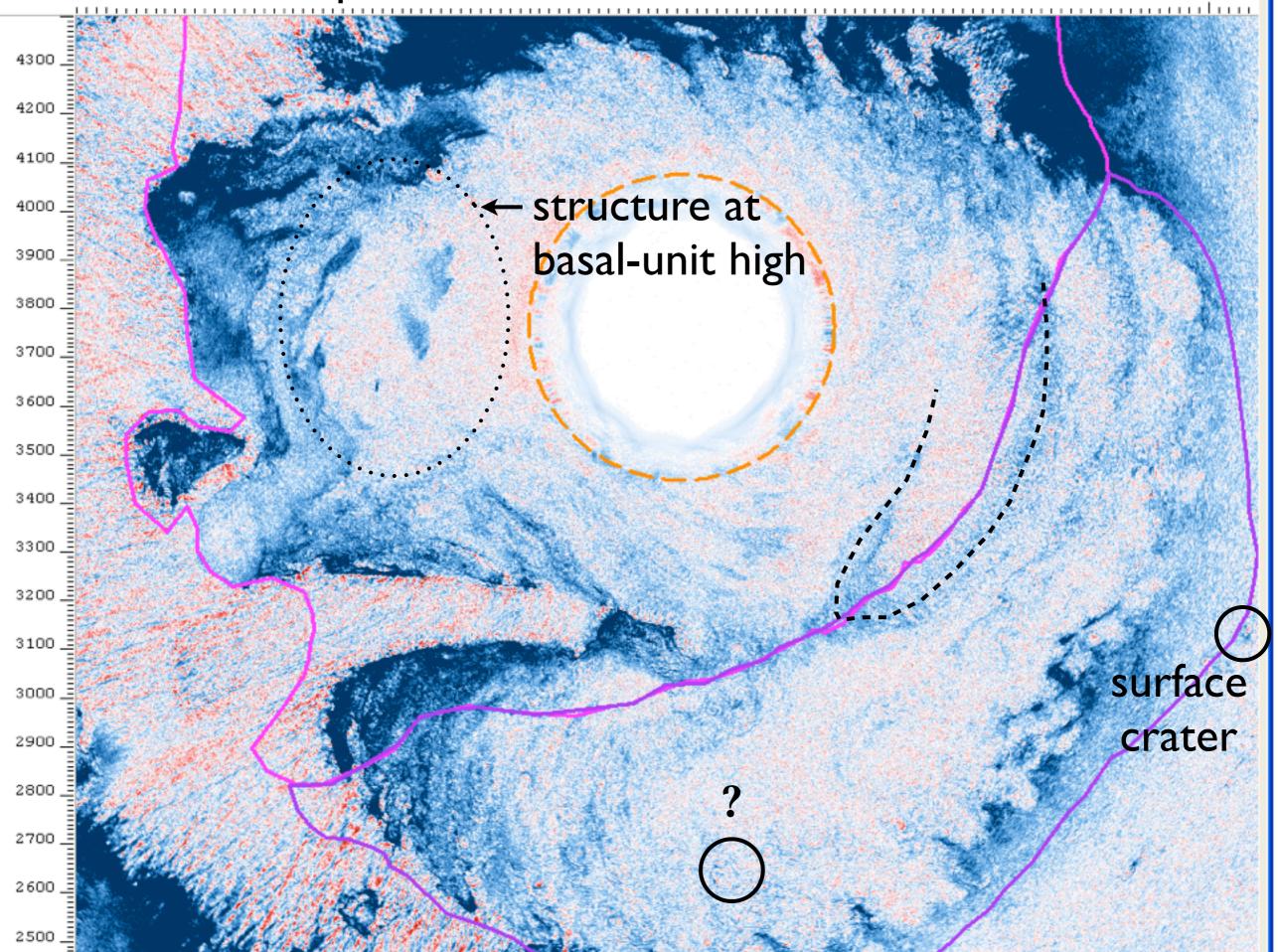


Expression of buried chasma evolves



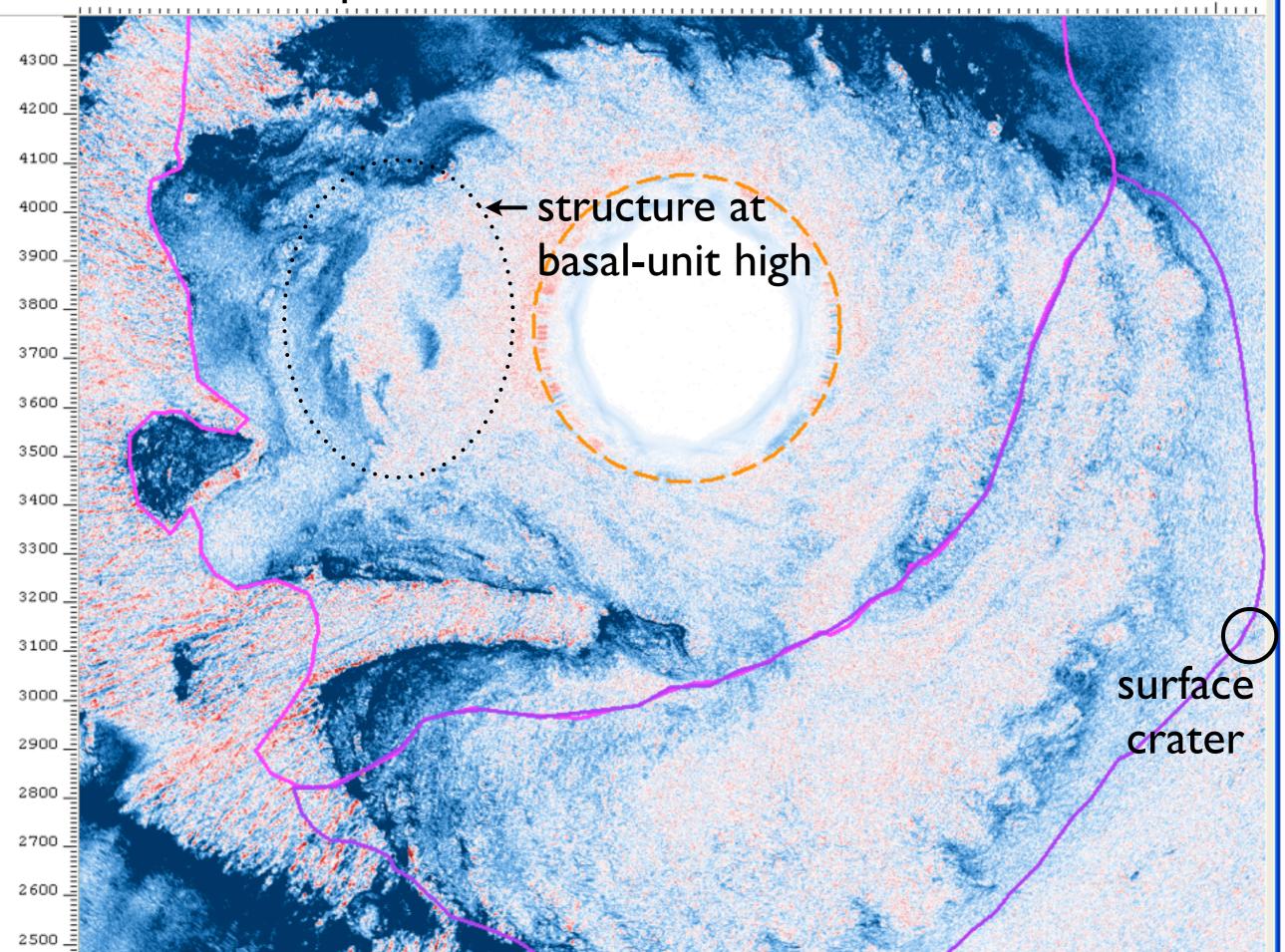
Expression of buried chasma evolves

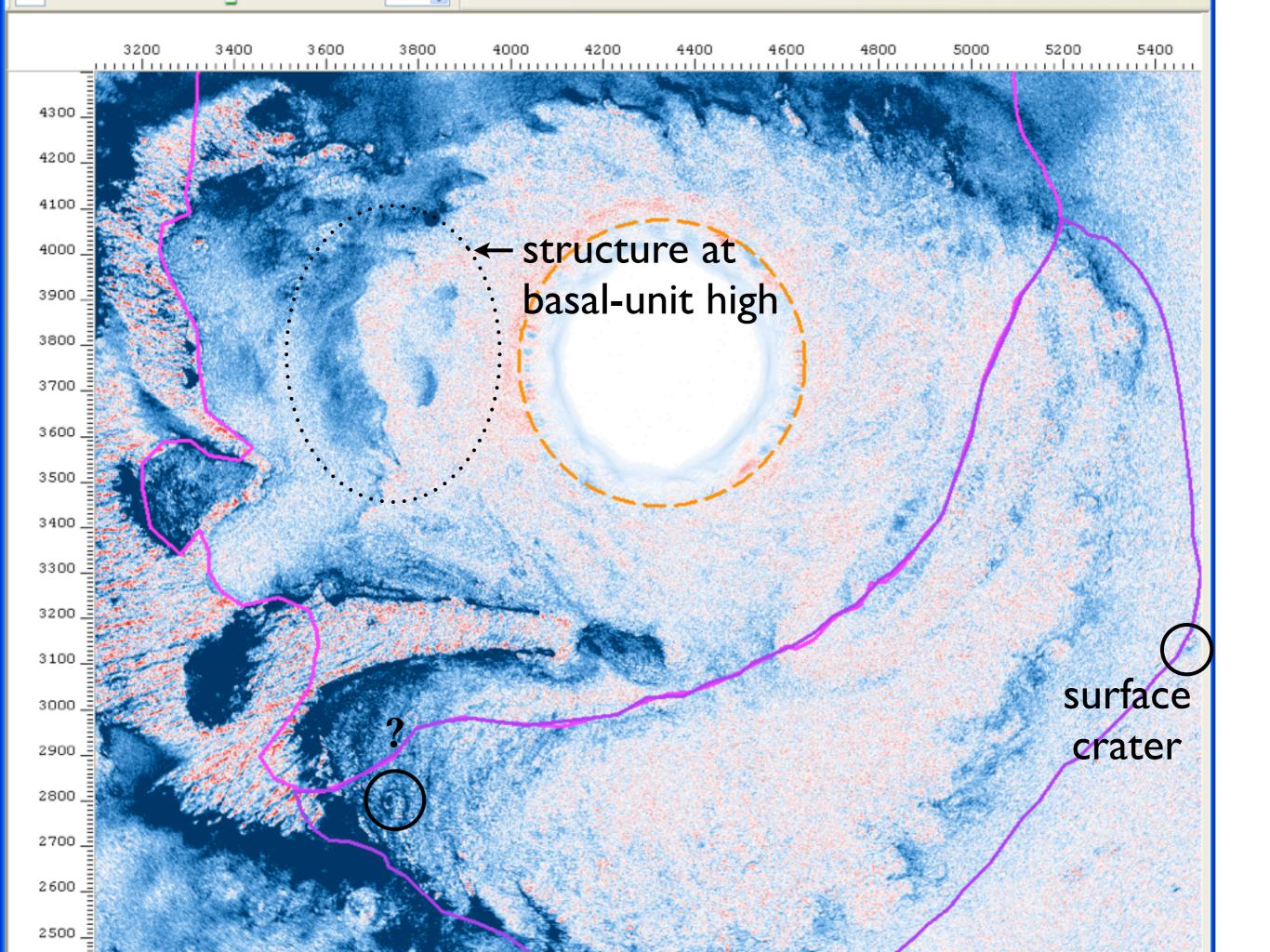
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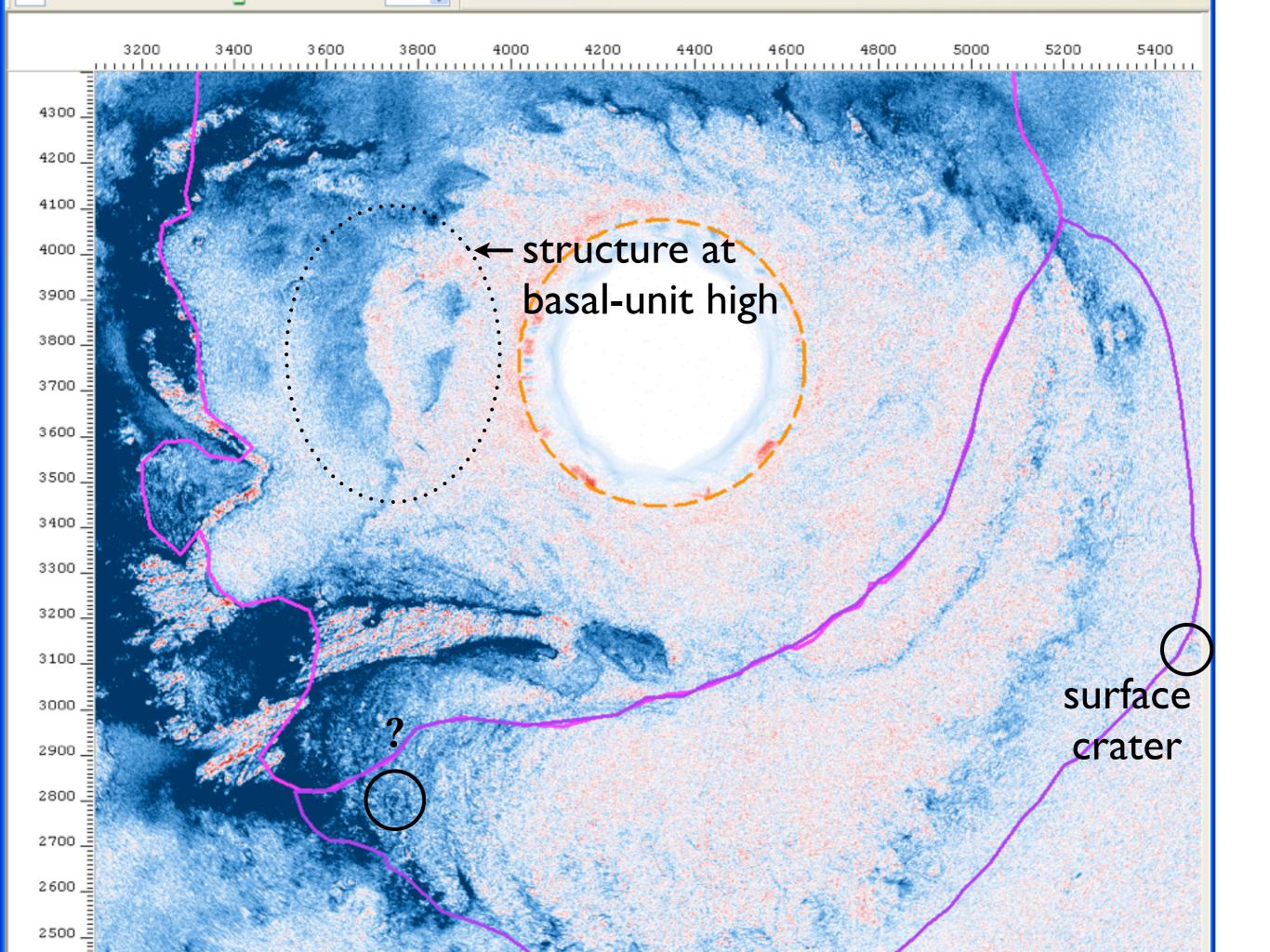


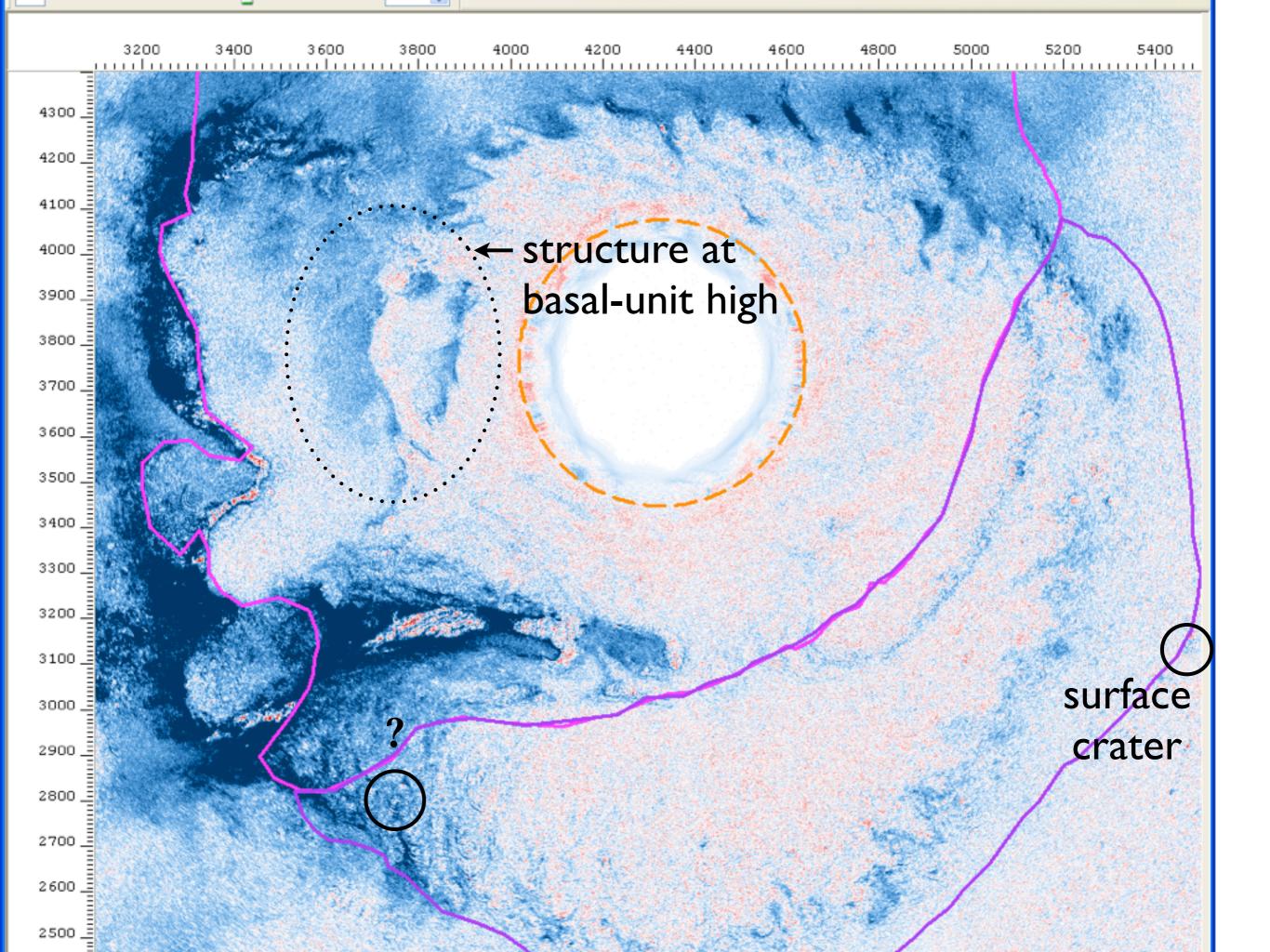
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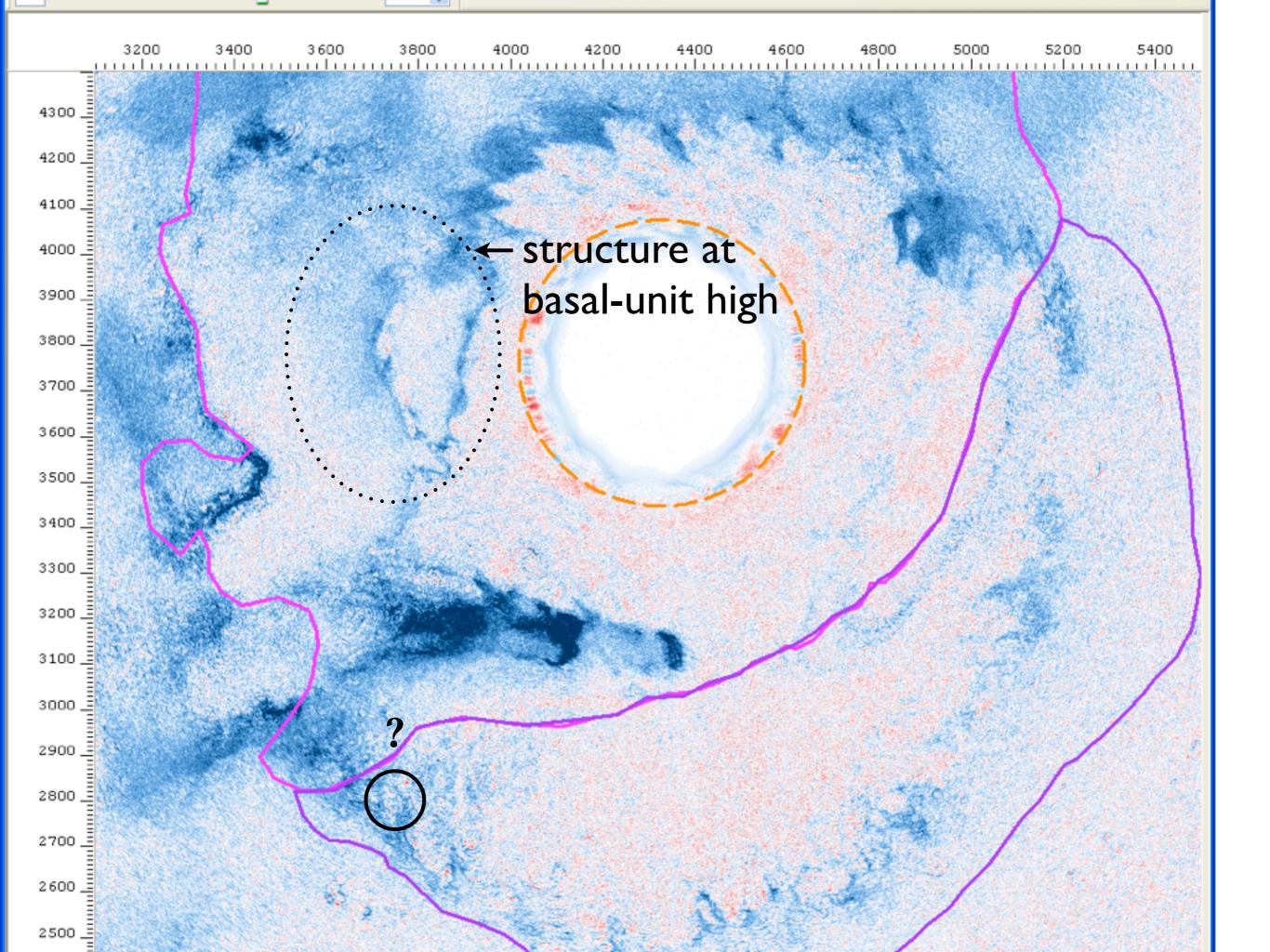
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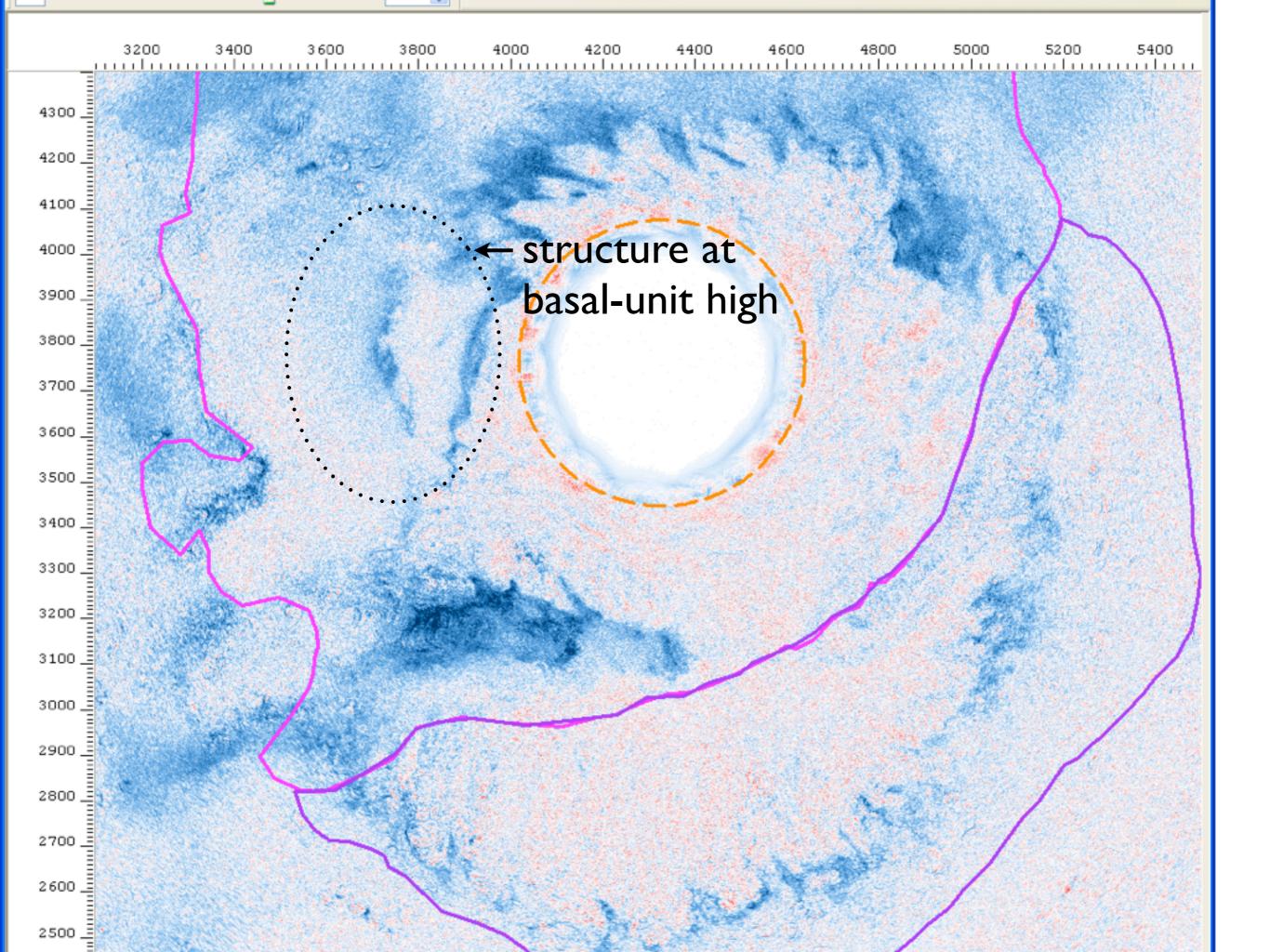


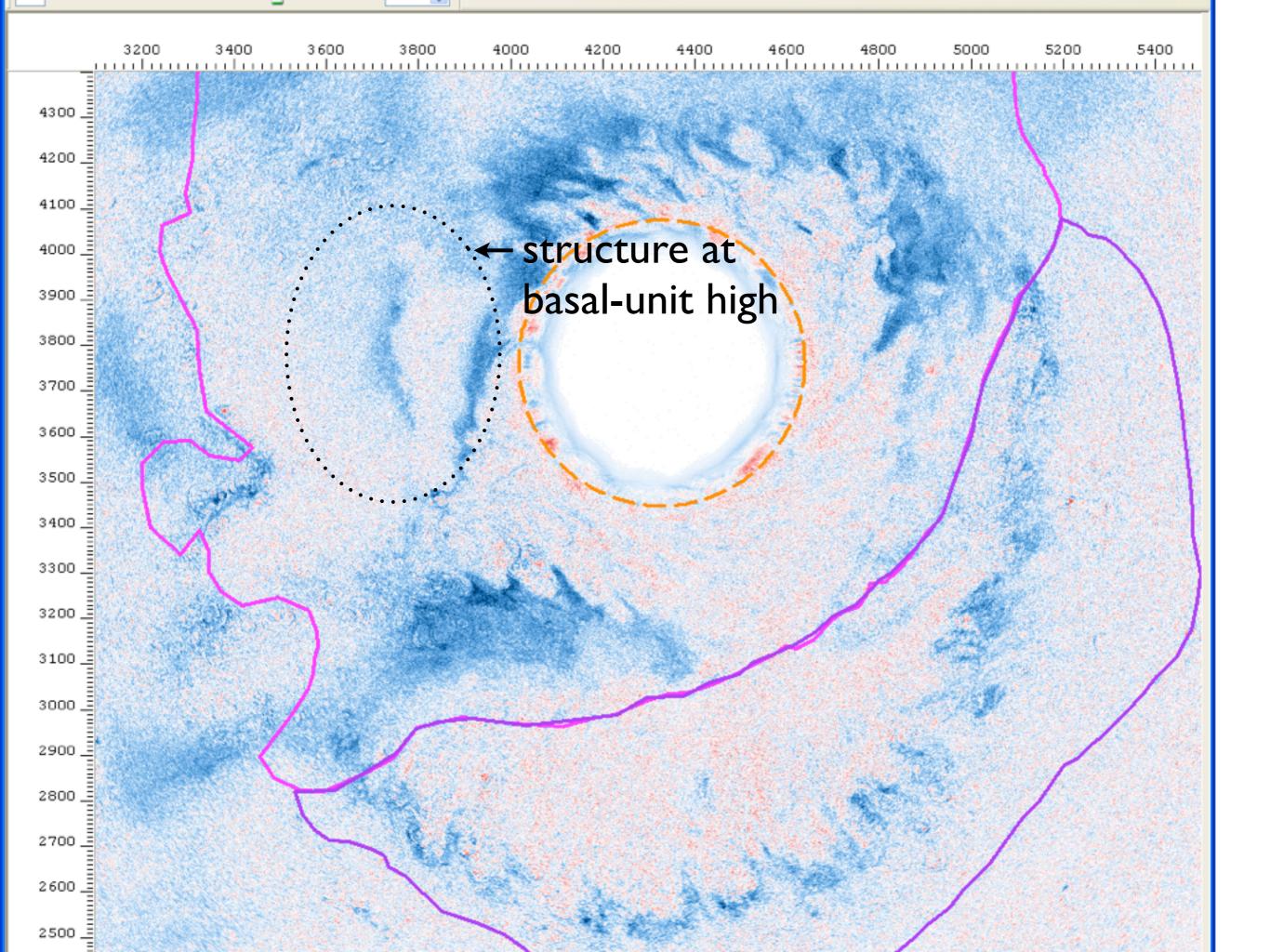


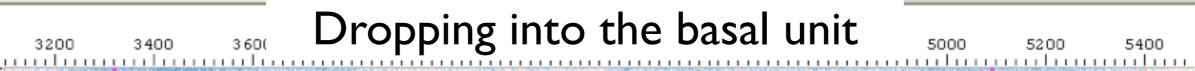




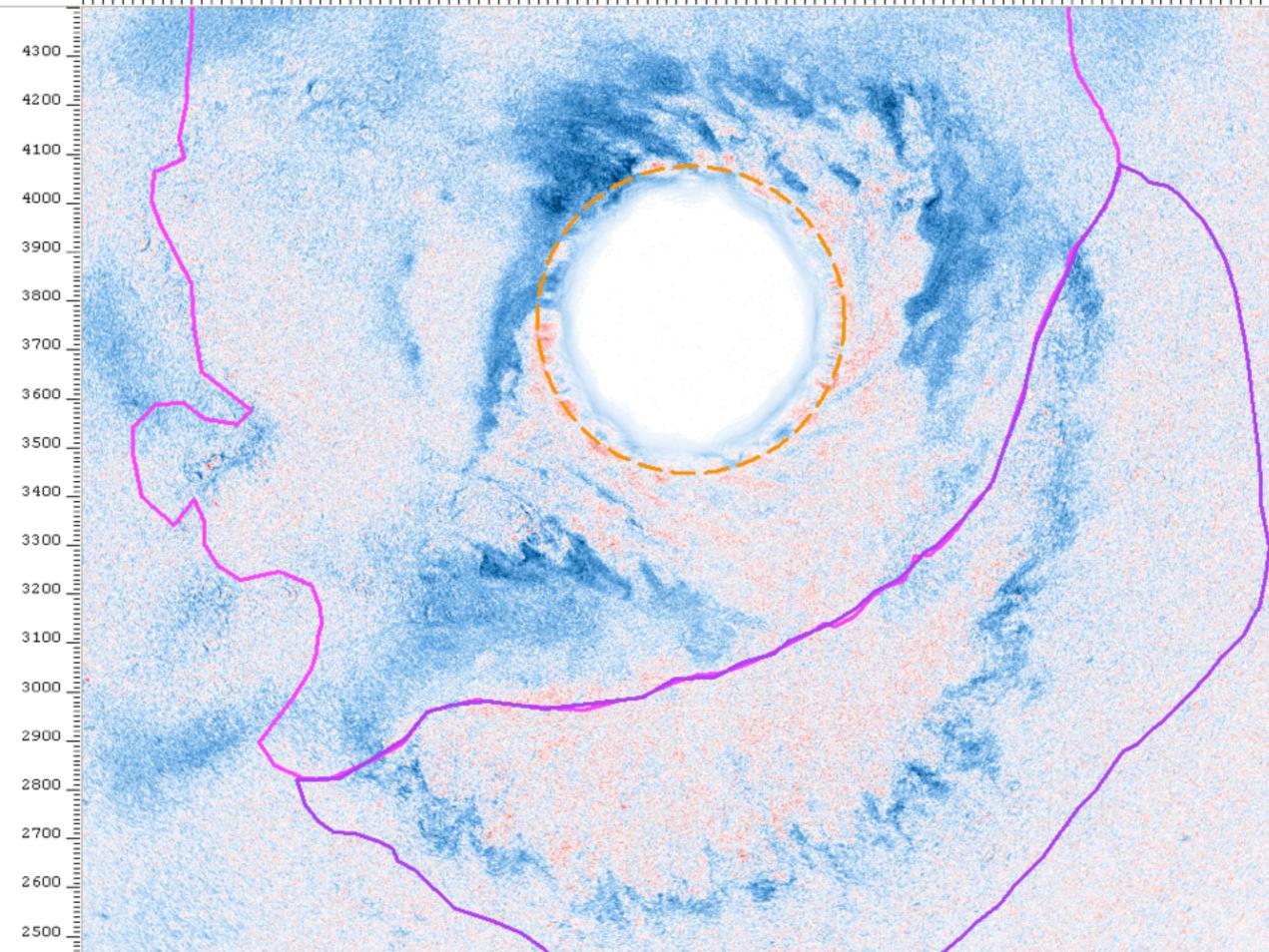




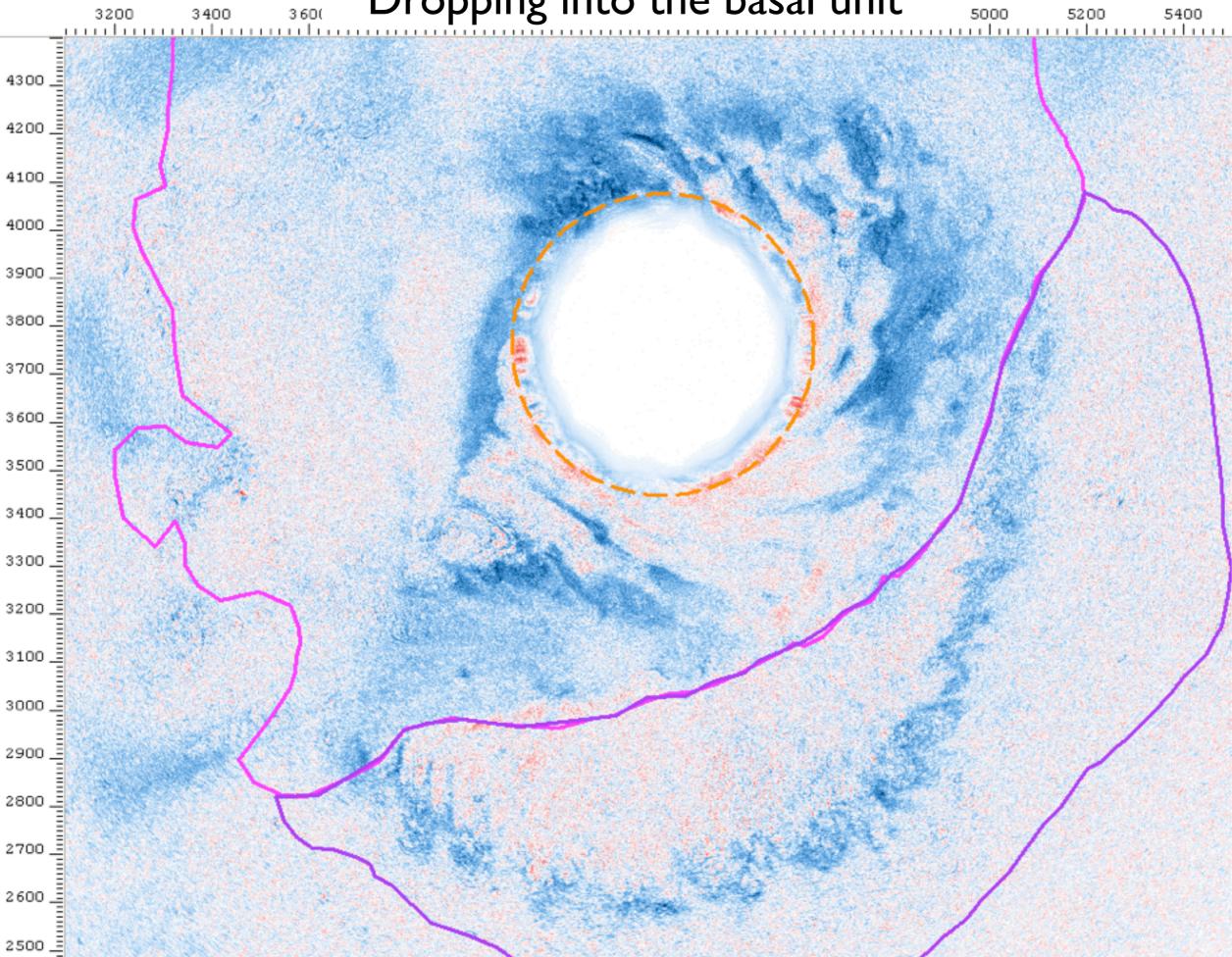


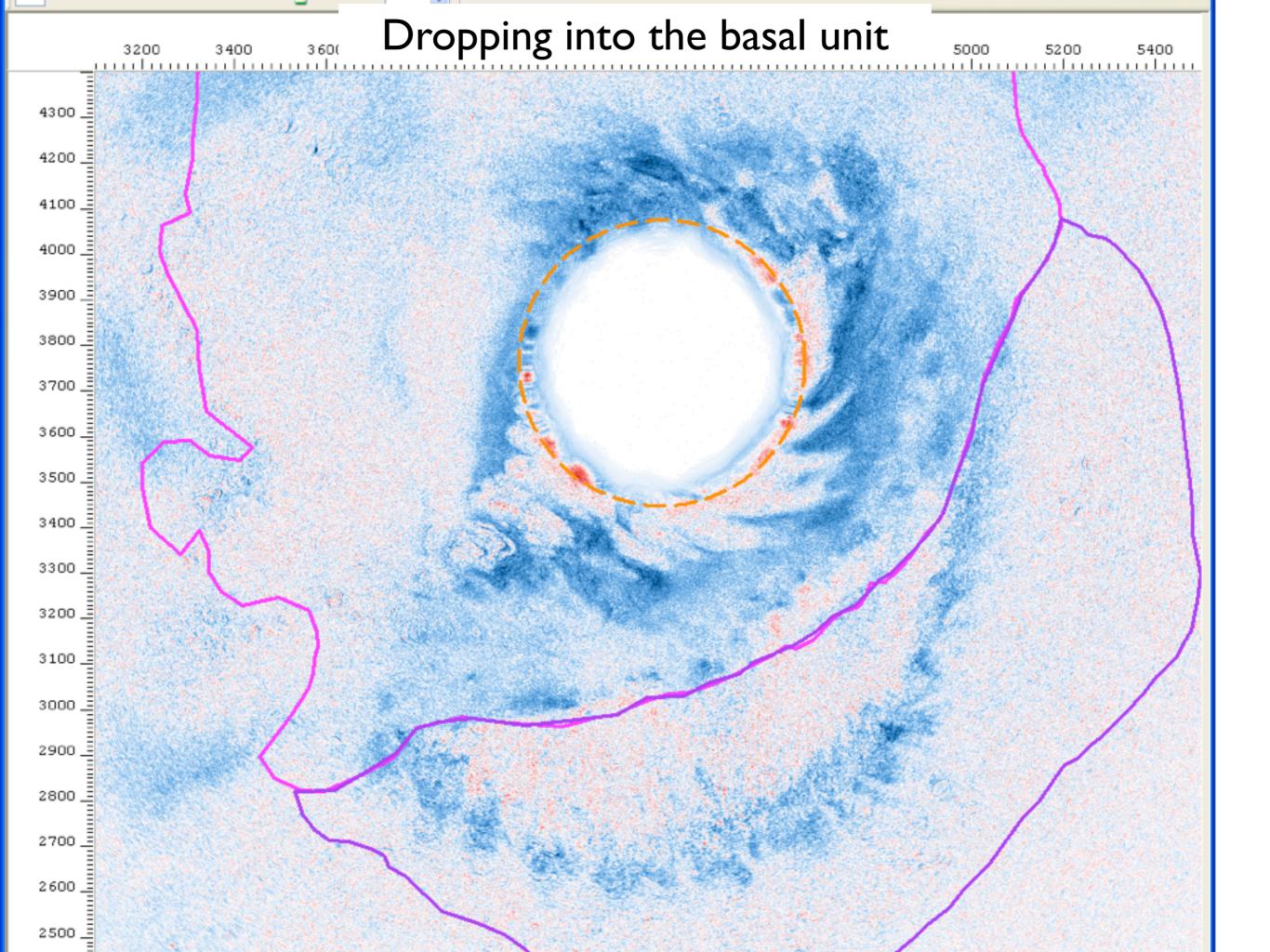


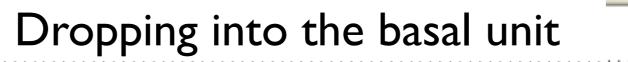




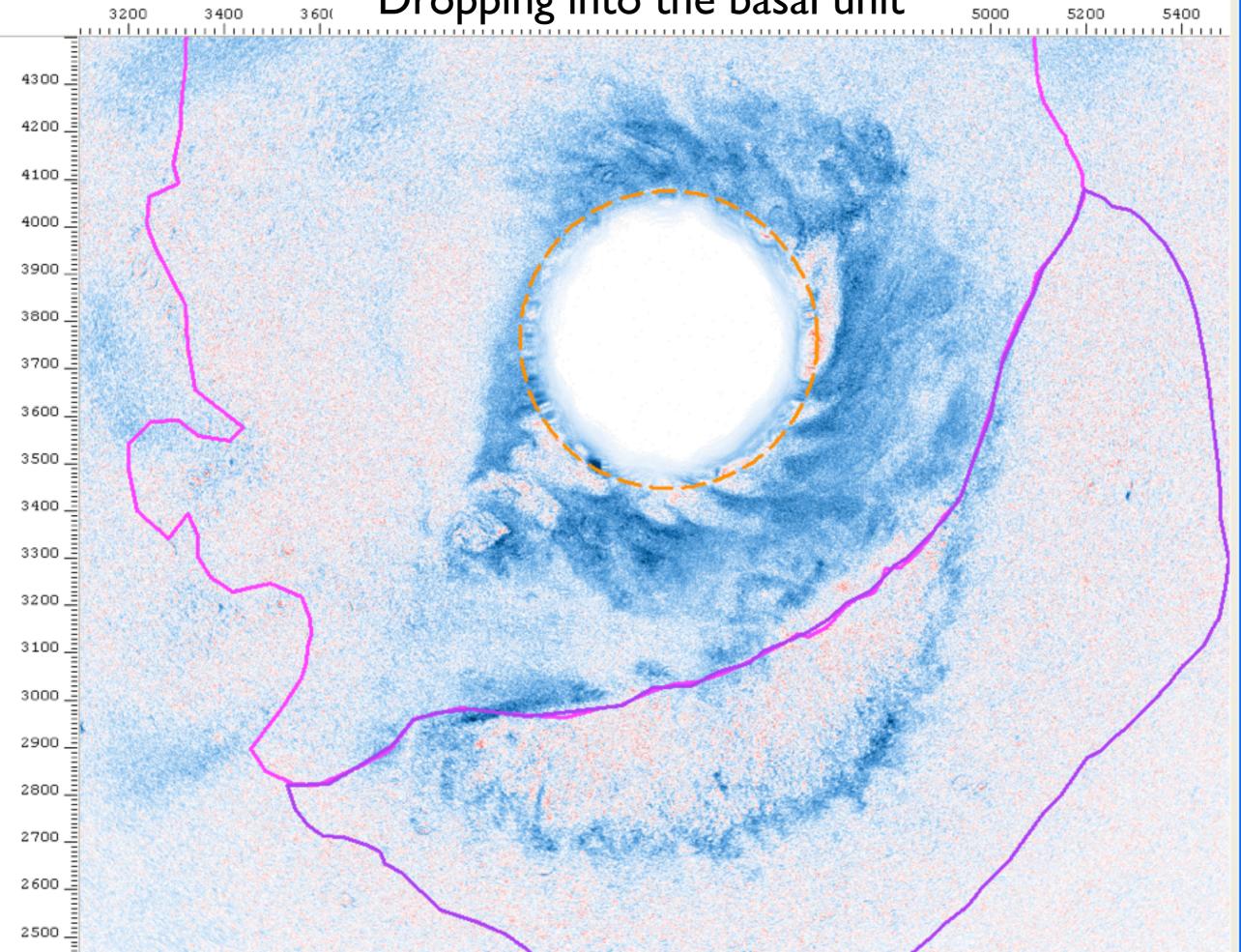
Dropping into the basal unit 5000 5200 5400

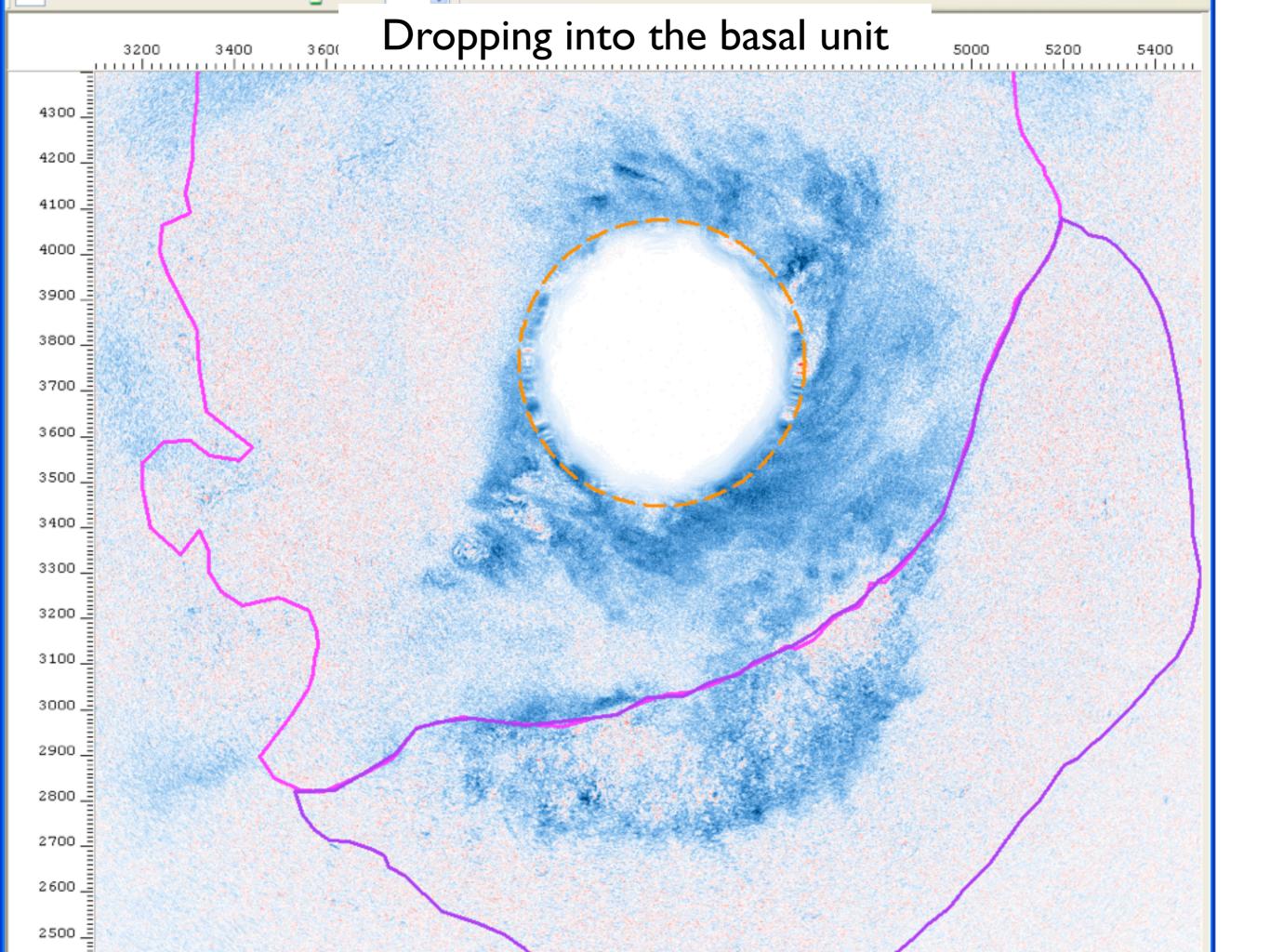


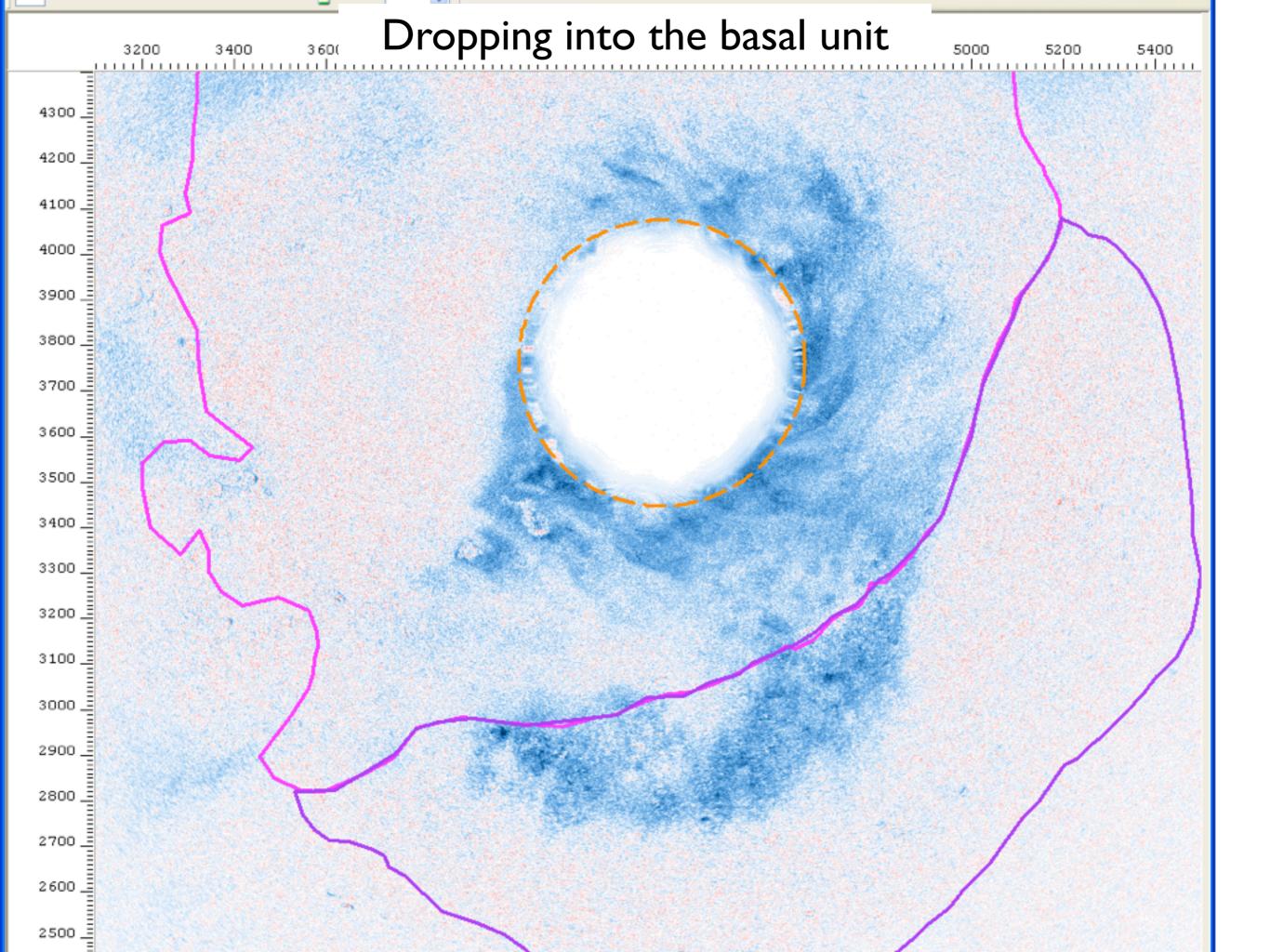


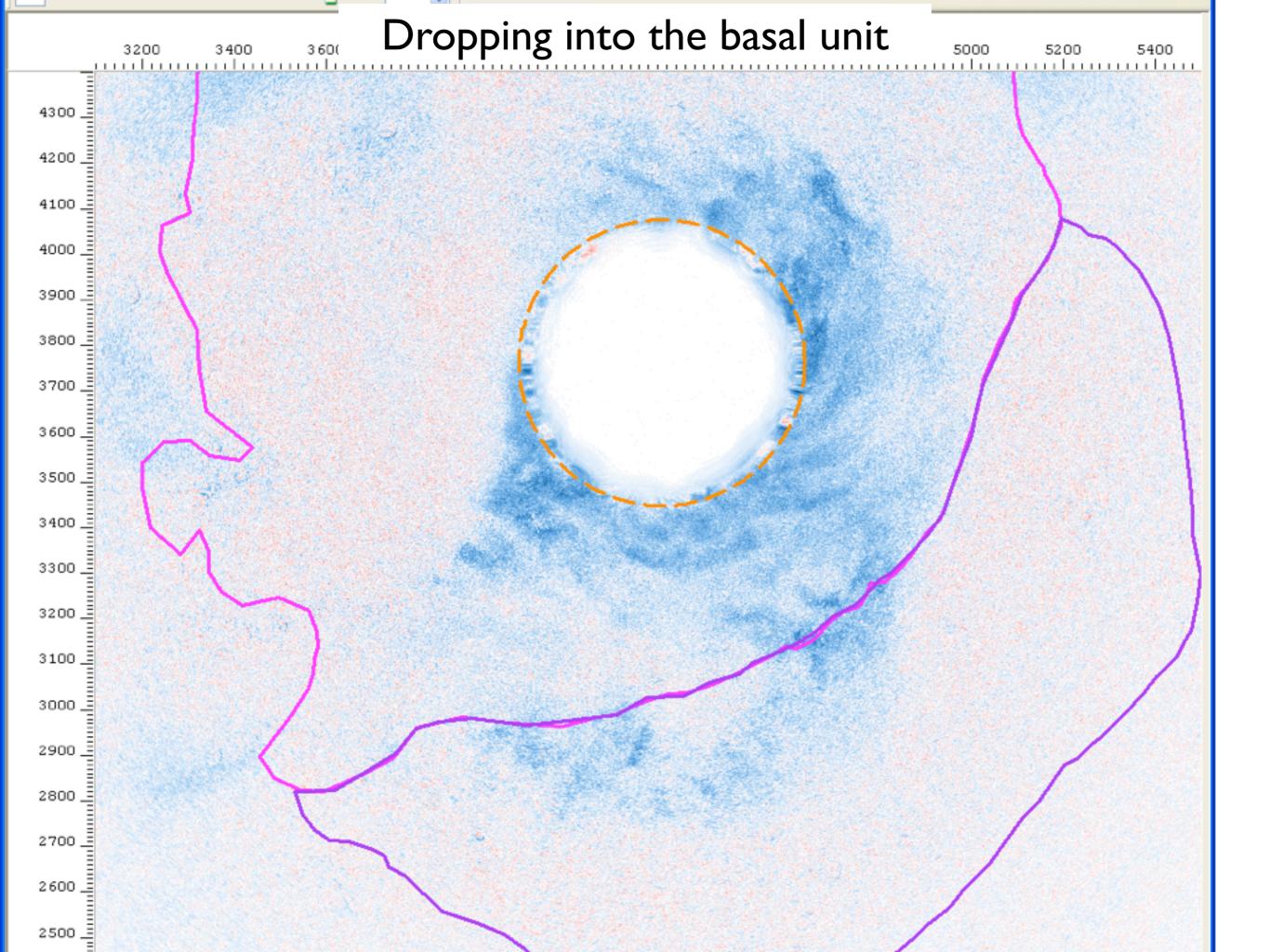


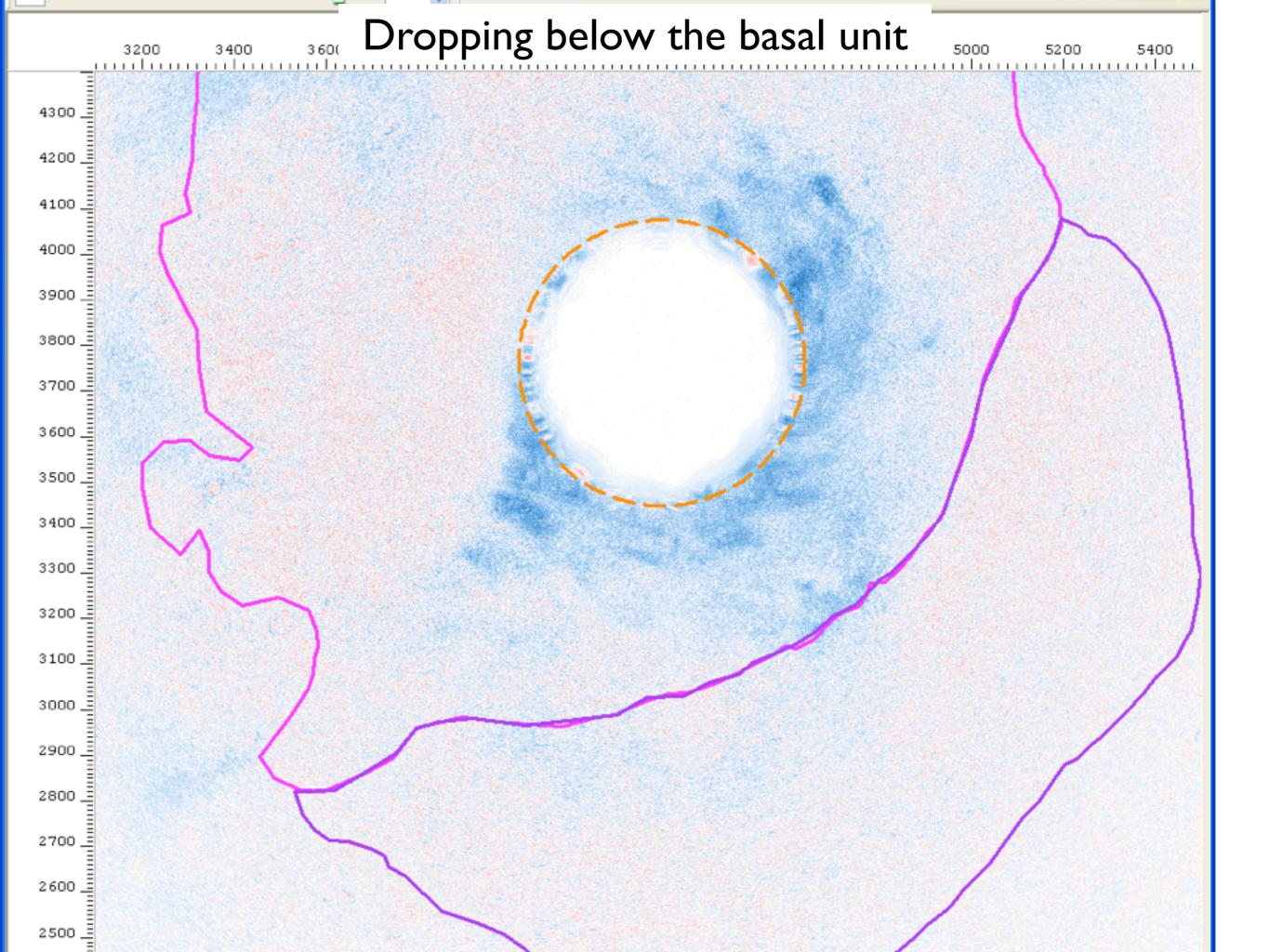


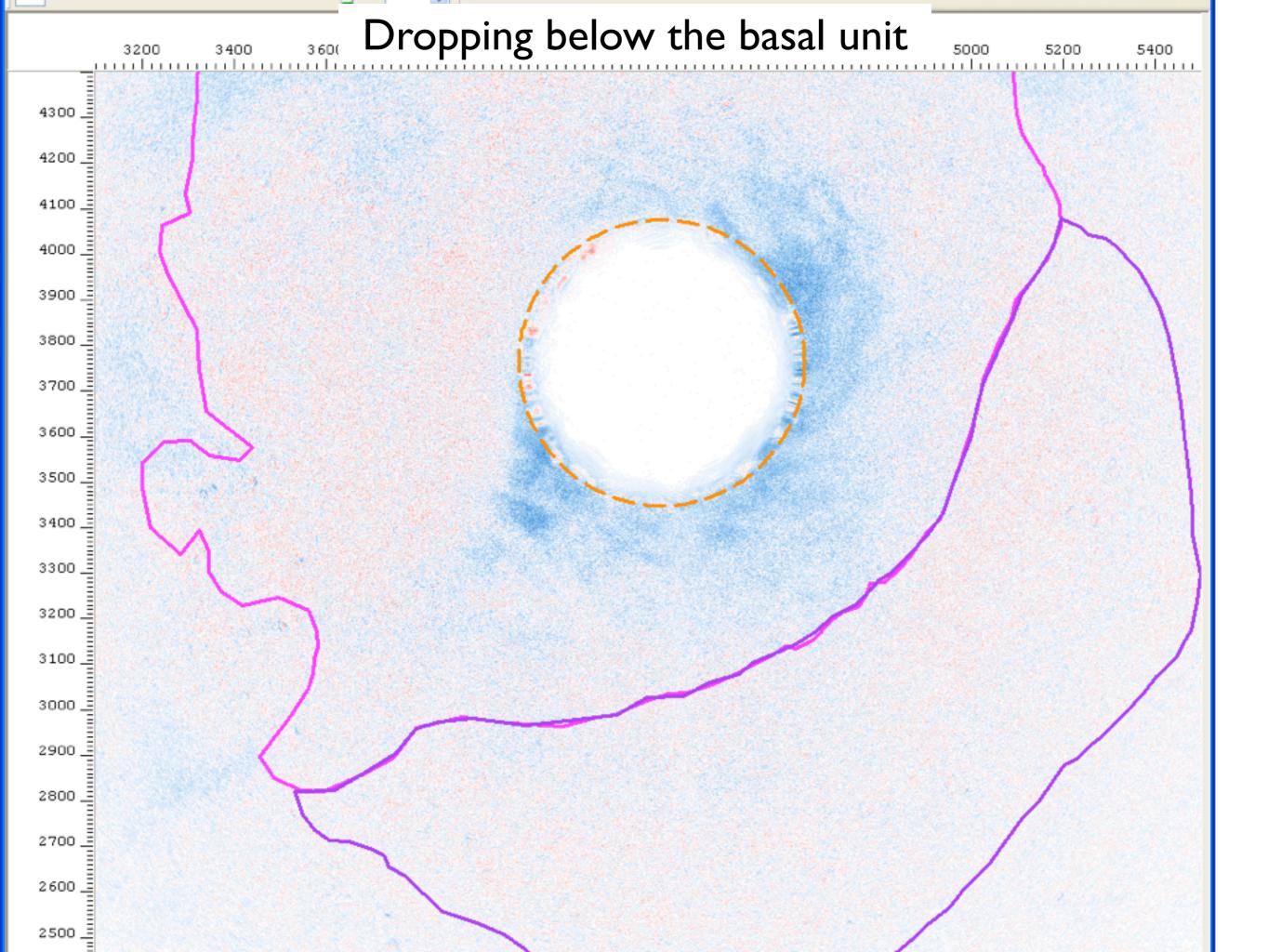


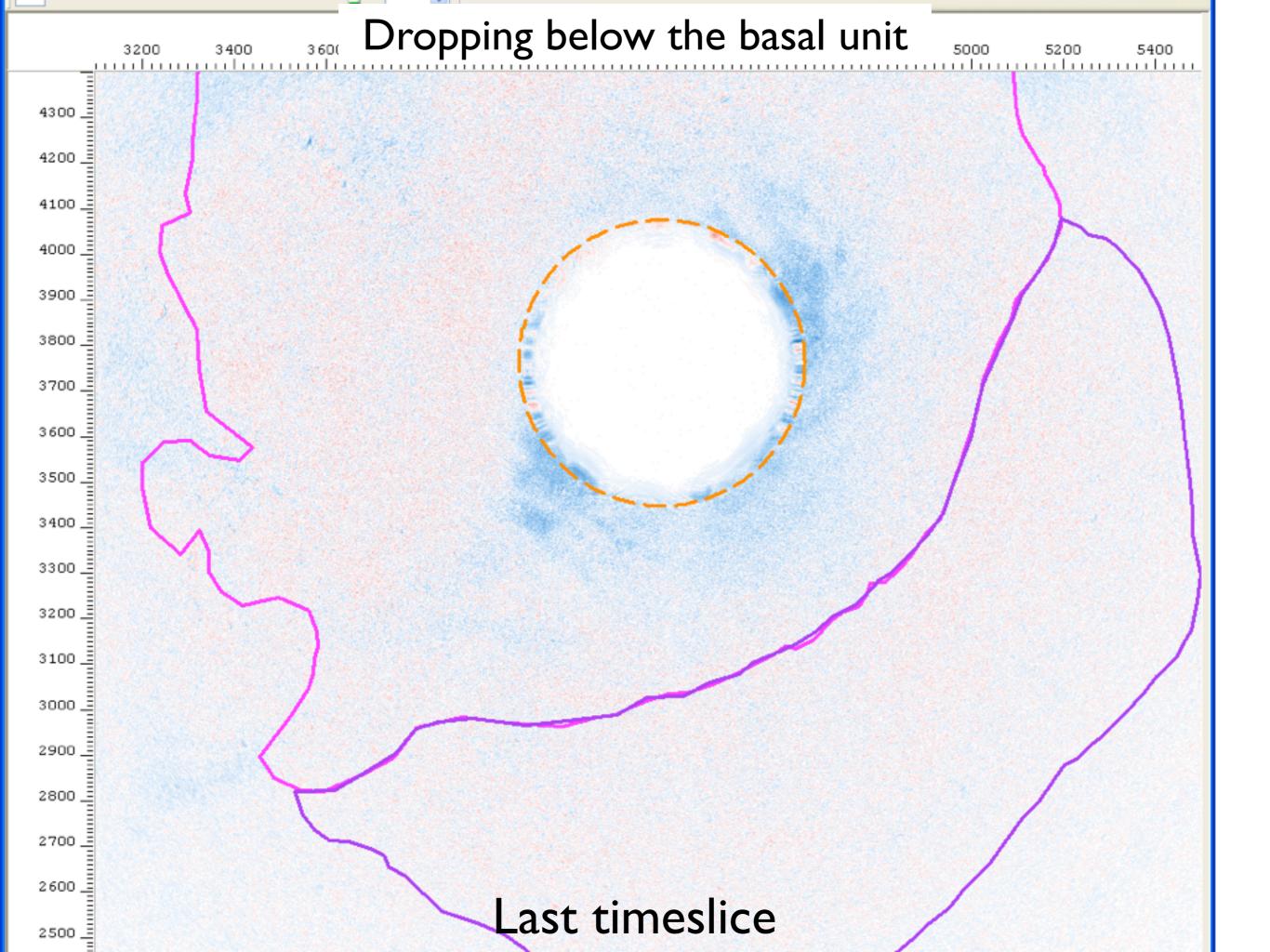








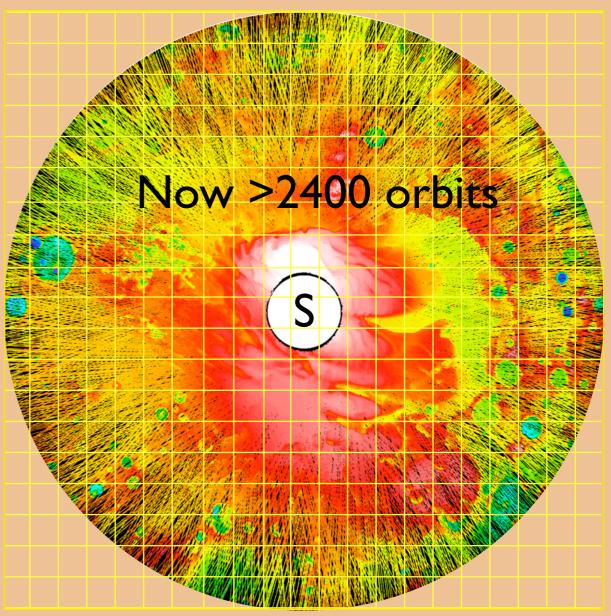




What's next?

- Planum Boreum 3D will be enhanced
 further, converted to
 depth, and released
 for public use.
- A Planum Australe
 3D will be coming
 in the near future!

Coverage 70–87°S (in MOLA elevation)



Conclusions

- 2D SHARAD analysis has yielded a wealth of discoveries in the polar ices.
- 3D binning and migration are adding a new level of clarity to the data, revealing intriguing new features.
- Results will shed new light on the nature & timing of the polar deposits.

Thanks to...

- SeisWare and Landmark for software access and support.
- NASA, ASI (Italian Space Agency), MRO Project, and SHARAD Team.
- NASA's Mars Data Analysis Program for funding both 3-D studies.